# ANALYSIS OF TRANSIENT ELECTROMAGENTIC SCATTERING FROM THREE DIMENSIONAL CAVITIES

PEIJUN LI\*, LI-LIAN WANG<sup>†</sup>, AND AIHUA WOOD<sup>‡</sup>

**Abstract.** This paper is concerned with the mathematical analysis of the time-domain Maxwell equations in a three-dimensional cavity. An exact transparent boundary condition (TBC) is developed to reformulate the open cavity scattering problem in an unbounded domain equivalently into an initial-boundary value problem in a bounded domain. The well-posedness and stability are studied for the reduced problem. Moreover, an *a priori* estimate is established for the electric field with a minimum regularity requirement for the data.

**Key words.** Time-domain Maxwell equations, transparent boundary conditions, three-dimensional cavity scattering problem, well-posedness, stability, *a priori* estimates

AMS subject classifications. 35Q61, 78A25, 78M30

1. Introduction. This paper is concerned with the mathematical analysis of an electromagnetic cavity scattering problem governed by the time-domain Maxwell equations. An open cavity can be broadly defined as a local perturbation of an infinite plane surface. The cavity scattering problems have significant industry and military applications such as the design of cavity-backed conformal antennas and the deliberate control in the form of enhancement or reduction of radar cross section.

The time-harmonic problems have been widely investigated by numerous researchers in the engineering and mathematical communities [2–4, 7, 22, 24, 30]. A large amount of information is available regarding their solutions for both the two-dimensional Helmholtz and the three-dimensional Maxwell equations including the overfilled cavity problems [7, 8, 27, 29, 41]. We refer to [23] for a good introduction to the cavity scattering problem. One may consult [14, 15, 31, 32] for recent accounts of finite element methods and integral equation methods for general electromagnetic scattering problems.

The time-domain problems have attracted much attention due to their capability of capturing wide-band signals and modeling more general material and nonlinearity [9,21,26,39,40,42]. However, rigorous mathematical analysis is very rare, especially for time-domain three-dimensional Maxwell equations. The transient cavity scattering problems have been mathematically studied in [19,20,34,36,37], where the focus was largely on (i) temporal discretization by e.g., the Newmark scheme; (ii) the reduction of the resulted system via frequency-domain TBC, and (iii) the analysis of the finite-element method for the reduced problem. To the best of our knowledge, the theoretical analysis of the time-domain Maxwell cavity problem itself was left undone and is still lacking.

In this work, we aim to analyze the transient electromagnetic wave scattering by an open cavity which is embedded in a perfectly conducting infinite ground plane.

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The cavity is filled with some inhomogeneous medium which may even protrude from its opening to above the ground plane, while the upper half space outside of the cavity is filled with a homogeneous medium. This problem is quite challenging due to the unbounded nature of the domain. In the frequency domain, various approaches have been investigated such as developing TBCs and designing perfectly matched layer techniques [1, 5, 10, 12, 17, 18, 25, 28, 29]. In contrast, much less study is devoted to address this issue in the time domain. In [11], the well-posedness and stability were shown for the electromagnetic obstacle scattering problem by using a time-domain TBC on the sphere. Here, we present a new time-domain transparent boundary condition on the hemisphere emcompassing the cavity and prove well-posedness and stability of the underlying scattering problem. The proofs are based on examining the well-posedness of the time-harmonic Maxwell equations with complex wavenumbers and applying the abstract inversion theorem of the Laplace transform. Moreover, an apriori estimate featured with explicit dependence on time and a minimum regularity requirement of the data, is established for the electric field by studying directly the time-domain Maxwell equations.

The outline of this paper is as follows. In section 2, we introduce the model problem and exploit the time-domain TBC to reduce the underlying problem to an initial-boundary value problem in a bounded domain. In section 3, we analyze two auxiliary problems pertinent to the reduced time-domain Maxwell equations to pave the way for the analysis of the main results in section 4. We study in section 4 the well-posedness and stability of the reduced Maxwell equations, and derive an a prior estimate with a minimum requirement for the regularity of the data. The paper is concluded with some remarks and directions for future work in section 5. To avoid distraction from the main results, we collect in the appendices some necessary notation and useful results on the Laplace transform, spherical harmonics, and functional spaces.

- 2. Formulation and reduction of the problem. In this section, we introduce the mathematical model of interest, and exploit the exact TBC to reduce the unbounded domain to a bounded one.
- **2.1.** A model problem. We describe the setting of a cavity problem and define some necessary notations. Denote by D the cavity embedded in the perfectly electrically conducting infinite ground plane  $\Gamma_g$ . Let  $S = \partial D \cap \mathbb{R}^3$ , the part of cavity walls below the ground, be Lipschitz continuous and perfectly electrically conducting. The cavity is filled with some inhomogeneous medium, which may protrude from its opening to above the ground plane. The medium is characterized by the dielectric permittivity  $\varepsilon$  and the magnetic permeability  $\mu$ , which satisfy

$$0 < \varepsilon_{\min} \le \varepsilon \le \varepsilon_{\max} < \infty, \quad 0 < \mu_{\min} \le \mu \le \mu_{\max} < \infty.$$

Here  $\varepsilon_{\min}$ ,  $\varepsilon_{\max}$ ,  $\mu_{\min}$ , and  $\mu_{\max}$  are constants. Therefore, the analysis in this work includes both planar and overfilled cavities. Let  $B_R^+$  and  $\Gamma_R^+$  be the half-ball and hemisphere above the ground plane, where the radius R is large enough to completely contain the possibly overfilled portion of the cavity. The unbounded region  $\mathbb{R}^3_+ \setminus \bar{B}_R^+$  is filled with some homogeneous material with a constant permittivity  $\varepsilon_0$  and a constant permeability  $\mu_0$ . Throughout this paper, we assume for simplicity in exposition that  $\varepsilon_0 = \mu_0 = 1$ . Finally, we denote by  $\Omega = B_R^+ \cup D$  the bounded domain in which our reduced initial-boundary value problem will be formulated.

It is evident that the problem geometry is not only applicable to the cavity problem but also to a broader class of scattering problems where the interface or the

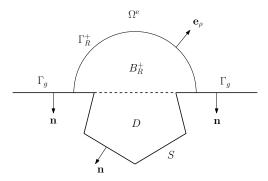


Fig. 2.1. A schematic diagram of the cavity problem geometry.

boundary is modeled as a local perturbation of an infinite plane.

Consider the system of time-domain Maxwell equations in  $\mathbb{R}^3_+ \cup D$  for t > 0:

(2.1) 
$$\begin{cases} \nabla \times \boldsymbol{E}(\boldsymbol{r},t) + \mu \partial_t \boldsymbol{H}(\boldsymbol{r},t) = \boldsymbol{0}, \\ \nabla \times \boldsymbol{H}(\boldsymbol{r},t) - \varepsilon \partial_t \boldsymbol{E}(\boldsymbol{r},t) = \boldsymbol{J}(\boldsymbol{r},t), \end{cases}$$

where  $r = (x, y, z) \in \mathbb{R}^3$ , E is the electric field, H is the magnetic field, and J is the electric current density which is assumed to be compactly supported in D. The system is constrained by the initial conditions:

(2.2) 
$$E|_{t=0} = E_0, \quad H|_{t=0} = H_0 \text{ in } \mathbb{R}^3_+ \cup D,$$

where  $E_0$  and  $H_0$  are also assumed to be compactly supported in D. We consider the perfectly electrical conducting boundary condition on the ground plane and cavity wall:

(2.3) 
$$\mathbf{n} \times \mathbf{E} = \mathbf{0}$$
 on  $\Gamma_g \cup S$ ,  $t > 0$ ,

where n is the unit outward normal vector on  $\Gamma_g \cup S$ . In addition, we impose the Silver-Müller radiation condition:

(2.4) 
$$\hat{\boldsymbol{r}} \times (\partial_t \boldsymbol{E} \times \hat{\boldsymbol{r}}) + \hat{\boldsymbol{r}} \times \partial_t \boldsymbol{H} = o(|\boldsymbol{r}|^{-1}), \text{ as } |\boldsymbol{r}| \to \infty, t > 0,$$

where  $\hat{\boldsymbol{r}} = \boldsymbol{r}/|\boldsymbol{r}|$ .

The purpose of this paper is to study the well-posedness and establish the stability for the time-domain electromagnetic cavity scattering problem (2.1)–(2.4). Hereafter, the expression " $a \lesssim b$ " stands for  $a \leq Cb$ , where C is a generic positive constant independent of any function and important parameters, which are clear from the context.

**2.2. Transparent boundary condition.** We introduce a TBC to reformulate the electromagnetic wave propagation problem from an open cavity into an equivalent initial-boundary value problem in a bounded domain. The essential idea is to design a boundary operator which maps the tangential component of the electric field to the tangential trace of the magnetic field.

More precisely, we consider the reduced initial-boundary value problem:

(2.5) 
$$\begin{cases} \nabla \times \boldsymbol{E} + \mu \partial_t \boldsymbol{H} = 0, \quad \nabla \times \boldsymbol{H} - \varepsilon \partial_t \boldsymbol{E} = \boldsymbol{J} & \text{in } \Omega, \ t > 0, \\ \boldsymbol{E}|_{t=0} = \boldsymbol{E}_0, \quad \boldsymbol{H}|_{t=0} = \boldsymbol{H}_0 & \text{in } \Omega, \\ \boldsymbol{n} \times \boldsymbol{E} = 0 & \text{on } S, \ t > 0, \\ \mathcal{T}[\boldsymbol{E}_{\Gamma_R^+}] = \boldsymbol{H} \times \boldsymbol{n} & \text{on } \Gamma_R^+, \ t > 0, \end{cases}$$

where  $E_{\Gamma_R^+}$  is the tangential trace of E on  $\Gamma_R^+$ , and  $\mathscr{T}$  is the time-domain electric-to-magentic (EtM) Calderon operator, as the counterpart of the time-harmonic setting with e.g., the spherical boundary (cf. [15,31]).

In what follows, we derive the formulation of the operator  $\mathscr{T}$  and analyze its important properties. Equivalently, we aim to prove the well-posedness and stability of the reduced problem (2.5). In particular, a priori estimates are established with minimum requirements for the regularity of the data. The proofs are based on the abstract inversion theorem of the Laplace transform and the a priori estimates for the time-harmonic Maxwell equations with a complex wavenumber. These will be the main topics of the forthcoming sections.

Since **J** is supported in D and  $\varepsilon_0 = \mu_0 = 1$ , the Maxwell equations (2.1) become

(2.6) 
$$\nabla \times \mathbf{E} + \partial_t \mathbf{H} = \mathbf{0}, \quad \nabla \times \mathbf{H} - \partial_t \mathbf{E} = \mathbf{0} \text{ in } \Omega^e := \mathbb{R}^3_+ \setminus \bar{B}^+_R, \quad t > 0.$$

Let  $\check{\boldsymbol{E}}(\boldsymbol{r},s) = \mathcal{L}(\boldsymbol{E})$  and  $\check{\boldsymbol{H}}(\boldsymbol{r},s) = \mathcal{L}(\boldsymbol{H})$  be the Laplace transforms of  $\boldsymbol{E}(\boldsymbol{r},t)$  and  $\boldsymbol{H}(\boldsymbol{r},t)$  with respect to t, respectively, where the complex variable

(2.7) 
$$s = s_1 + is_2 \text{ with } s_1, s_2 \in \mathbb{R}, s_1 > 0, i = \sqrt{-1},$$

as seen in Appendix A. Recall that

$$\mathscr{L}(\partial_t \mathbf{E}) = s \mathbf{\breve{E}} - \mathbf{E}_0, \quad \mathscr{L}(\partial_t \mathbf{H}) = s \mathbf{\breve{H}} - \mathbf{H}_0.$$

Taking the Laplace transform of (2.6), and noting that  $E_0$ ,  $H_0$  are supported in D, we obtain the time-harmonic Maxwell equations with complex parameters:

(2.8) 
$$\nabla \times \mathbf{E} + s\mathbf{H} = \mathbf{0}, \quad \nabla \times \mathbf{H} - s\mathbf{E} = \mathbf{0} \quad \text{in } \Omega^e, \quad s_1 > 0.$$

Let  $h_n^{(1)}(z)$  be the spherical Hankel function of the first kind of order n (cf. [38]). We introduce the vector wave functions

(2.9) 
$$\begin{cases} \boldsymbol{M}_{n}^{m}(\rho,\theta,\varphi) = \nabla \times (\boldsymbol{r}h_{n}^{(1)}(\mathrm{i}s\rho)X_{n}^{m}(\theta,\varphi)), \\ \boldsymbol{N}_{n}^{m}(\rho,\theta,\varphi) = -s^{-1}\nabla \times \boldsymbol{M}_{n}^{m}(\theta,\varphi), \end{cases}$$

where  $(\rho, \theta, \varphi)$  is the spherical coordinates, and  $X_n^m$  is the rescaled spherical harmonic function defined in (B.2). We refer to Appendix B for the properties of spherical harmonics and the relevant calculus to be used throughout this paper.

The vector wave functions in (2.9) are the radiation solutions of (2.8) in  $\mathbb{R}^3 \setminus \{0\}$  (cf. [31, Thm. 9.16]):

(2.10) 
$$\nabla \times \boldsymbol{M}_{n}^{m}(\rho, \theta, \varphi) + s\boldsymbol{N}_{n}^{m}(\rho, \theta, \varphi) = \boldsymbol{0}, \quad \nabla \times \boldsymbol{N}_{n}^{m}(\rho, \theta, \varphi) - s\boldsymbol{M}_{n}^{m}(\rho, \theta, \varphi) = \boldsymbol{0}.$$

It can be verified from (2.9) that the vector wave functions satisfy

(2.11) 
$$\boldsymbol{M}_{n}^{m}(\rho,\theta,\varphi) = h_{n}^{(1)}(\mathrm{i}s\rho)\nabla_{\Gamma}X_{n}^{m}(\theta,\varphi) \times \boldsymbol{e}_{\rho}$$

and

$$\boldsymbol{N}_{n}^{m}(\rho,\theta,\varphi) = -\frac{\sqrt{n(n+1)}}{s\rho} \left( h_{n}^{(1)}(\mathrm{i}s\rho) + \mathrm{i}s\rho(h_{n}^{(1)})'(\mathrm{i}s\rho) \right) \boldsymbol{X}_{n}^{m}(\theta,\varphi)$$

$$-\frac{n(n+1)}{s\rho} h_{n}^{(1)}(\mathrm{i}s\rho) Y_{n}^{m}(\theta,\varphi) \boldsymbol{e}_{\rho}.$$
(2.12)

Once again, we refer to Appendix B for the notation and definition. A simple calculation yields

$$\begin{cases} \boldsymbol{e}_{\theta} \times \boldsymbol{M}_{n}^{m}(\rho, \pi/2, \varphi) = 0 & \text{for } |m| \leq n, \ m+n = \text{even}, \ n \in \mathbb{N}, \\ \boldsymbol{e}_{\theta} \times \boldsymbol{N}_{n}^{m}(\rho, \pi/2, \varphi) = 0 & \text{for } |m| \leq n, \ m+n = \text{odd}, \ n \in \mathbb{N}. \end{cases}$$

Therefore, in the domain  $\Omega^{e}$ , the solution of the electric field  $\check{\boldsymbol{E}}(\rho,\theta,\varphi,s)$ , which satisfies the perfectly electric conducting condition  $\boldsymbol{n}\times\check{\boldsymbol{E}}=0$  on  $\Gamma_{g}$ , i.e.,  $\boldsymbol{e}_{\theta}\times\check{\boldsymbol{E}}(\rho,\pi/2,\varphi,s)=0$ , can be written in the series:

(2.13) 
$$\breve{\boldsymbol{E}}(\rho, \theta, \varphi, s) = \sum_{|m| \le n}^{\text{odd}} \alpha_n^m(s) \, \boldsymbol{N}_n^m(\rho, \theta, \varphi) + \sum_{|m| \le n}^{\text{even}} \beta_n^m(s) \, \boldsymbol{M}_n^m(\rho, \theta, \varphi),$$

which is uniformly convergent on compact subsets in  $\Omega^e$ . The corresponding magnetic field  $\check{\boldsymbol{H}}$  is given by

(2.14) 
$$\check{\boldsymbol{H}} = -s^{-1}\nabla \times \check{\boldsymbol{E}} = -\sum_{|m| \le n}^{\text{odd}} \alpha_n^m(s) \boldsymbol{M}_n^m(\rho, \theta, \varphi) + \sum_{|m| \le n}^{\text{even}} \beta_n^m(s) \boldsymbol{N}_n^m(\rho, \theta, \varphi).$$

Note that in the above, we compressed the summation notation as in Appendix B.

To deduce the explicit representation of the EtM Calderon operator, we need to express  $\boldsymbol{E}_{\Gamma_R^+} = -\boldsymbol{e}_\rho \times (\boldsymbol{e}_\rho \times \boldsymbol{E})$  and  $\boldsymbol{H} \times \boldsymbol{e}_\rho$  on  $\Gamma_R^+$  in terms of the coefficients  $\alpha_n^m$  and  $\beta_n^m$ . From the definition (2.11), one verifies that

$$\begin{split} &-\boldsymbol{e}_{\rho}\times(\boldsymbol{e}_{\rho}\times\boldsymbol{M}_{n}^{m}(\rho,\theta,\varphi))=-\sqrt{n(n+1)}\,h_{n}^{(1)}(\mathrm{i}s\rho)\boldsymbol{Y}_{n}^{m}(\theta,\varphi),\\ &-\boldsymbol{e}_{\rho}\times(\boldsymbol{e}_{\rho}\times\boldsymbol{N}_{n}^{m}(\rho,\theta,\varphi))=-\frac{\sqrt{n(n+1)}}{s\rho}\,\left(h_{n}^{(1)}(\mathrm{i}s\rho)+\mathrm{i}s\rho\,h_{n}^{(1)}{}'(\mathrm{i}s\rho)\right)\boldsymbol{X}_{n}^{m}(\theta,\varphi), \end{split}$$

and

$$\begin{split} & \boldsymbol{e}_{\rho} \times \boldsymbol{M}_{n}^{m}(\rho, \theta, \varphi) = \sqrt{n(n+1)} \, h_{n}^{(1)}(\mathrm{i} s \rho) \boldsymbol{X}_{n}^{m}(\theta, \varphi), \\ & \boldsymbol{e}_{\rho} \times \boldsymbol{N}_{n}^{m}(\rho, \theta, \varphi) = -\frac{\sqrt{n(n+1)}}{s \rho} \, \left( h_{n}^{(1)}(\mathrm{i} s \rho) + \mathrm{i} s \rho \, h_{n}^{(1)'}(\mathrm{i} s \rho) \right) \boldsymbol{Y}_{n}^{m}(\theta, \varphi). \end{split}$$

Therefore, by (2.13), the tangential component of the electric field along  $\Gamma_R^+$  is

$$\check{\boldsymbol{E}}_{\Gamma_{R}^{+}} = -\sum_{|m| \leq n}^{\text{odd}} \frac{\sqrt{n(n+1)}}{sR} \left( h_{n}^{(1)}(\mathbf{i}sR) + \mathbf{i}sR \, h_{n}^{(1)'}(\mathbf{i}sR) \right) \alpha_{n}^{m}(s) \, \boldsymbol{X}_{n}^{m}(\theta, \varphi) 
+ \sum_{|m| \leq n}^{\text{even}} \sqrt{n(n+1)} \, h_{n}^{(1)}(\mathbf{i}sR) \beta_{n}^{m}(s) \, \boldsymbol{Y}_{n}^{m}(\theta, \varphi),$$

and similarly, by (2.14), the tangential trace of the magnetic field along  $\Gamma_R^+$  is

$$\check{\boldsymbol{H}} \times \boldsymbol{e}_{\rho} = \sum_{|m| \leq n}^{\text{odd}} \sqrt{n(n+1)} h_n^{(1)}(\mathrm{i}sR) \alpha_n^m(s) \boldsymbol{X}_n^m(\theta, \varphi) 
+ \sum_{|m| \leq n}^{\text{even}} \frac{\sqrt{n(n+1)}}{sR} \left( h_n^{(1)}(\mathrm{i}sR) + \mathrm{i}sR h_n^{(1)}'(\mathrm{i}sR) \right) \beta_n^m(s) \boldsymbol{Y}_n^m(\theta, \varphi).$$

Consequently, we have the following explicit representation of the frequency domain EtM Calderon operator  $\mathcal{B}$ : given any tangential component of the electric field along  $\Gamma_R^+$  with the expansion:

(2.15) 
$$\boldsymbol{u} = \sum_{|m| \le n}^{\text{odd}} \alpha_n^m \boldsymbol{X}_n^m + \sum_{|m| \le n}^{\text{even}} \beta_n^m \boldsymbol{Y}_n^m,$$

the tangential trace of the magnetic field along  $\Gamma_R^+$  is

$$(2.16) \qquad \mathscr{B}[\boldsymbol{u}] = -\sum_{|m| \le n}^{\text{odd}} \frac{sR}{1 + r_n^{(1)}(\mathrm{i}sR)} \, \alpha_n^m \, \boldsymbol{X}_n^m - \sum_{|m| \le n}^{\text{even}} \frac{(1 + r_n^{(1)}(\mathrm{i}sR))}{sR} \, \beta_n^m \, \boldsymbol{Y}_n^m,$$

where

(2.17) 
$$r_n^{(1)}(z) = \frac{z h_n^{(1)'}(z)}{h_n^{(1)}(z)}.$$

We now analyze some properties of the EtM Calderon operator, and refer to Appendix D for the definitions of the function spaces to be used hereafter.

LEMMA 2.1. The Calderon operator  $\mathscr{B}: \mathbf{H}^{-1/2}(\operatorname{curl}, \Gamma_R^+) \to \mathbf{H}^{-1/2}(\operatorname{div}, \Gamma_R^+)$  is continuous.

*Proof.* For any  $u, v \in H^{-1/2}(\text{curl}, \Gamma_R^+)$ , we can expand

$$\boldsymbol{u} = \sum_{|m| < n}^{\text{odd}} u_{1n}^m \boldsymbol{X}_n^m + \sum_{|m| < n}^{\text{even}} u_{2n}^m \boldsymbol{Y}_n^m, \quad \boldsymbol{v} = \sum_{|m| < n}^{\text{odd}} v_{1n}^m \boldsymbol{X}_n^m + \sum_{|m| < n}^{\text{even}} v_{2n}^m \boldsymbol{Y}_n^m.$$

Then we have from (2.16) that

$$\mathscr{B}[\boldsymbol{u}] = -\sum_{|m| \le n}^{\text{odd}} \frac{sR \, u_{1n}^m}{1 + r_n^{(1)}(\mathrm{i} sR)} \, \boldsymbol{X}_n^m - \sum_{|m| \le n}^{\text{even}} \frac{(1 + r_n^{(1)}(\mathrm{i} sR)) u_{2n}^m}{sR} \, \boldsymbol{Y}_n^m.$$

To prove the lemma, it is required to estimate

$$(2.18) \quad \langle \mathscr{B}[\boldsymbol{u}], \boldsymbol{v} \rangle = -\sum_{|m| < n}^{\text{odd}} \frac{sR}{1 + r_n^{(1)}(\mathrm{i}sR)} u_{1n}^m \, \bar{v}_{1n}^m - \sum_{|m| < n}^{\text{even}} \frac{(1 + r_n^{(1)}(\mathrm{i}sR))}{sR} u_{2n}^m \, \bar{v}_{2n}^m.$$

It follows from the Cauchy-Schwarz inequality that

$$\begin{split} |\langle \mathscr{B}[\boldsymbol{u}], \boldsymbol{v} \rangle| &\leq \left[ \sum_{|m| \leq n}^{\text{odd}} \frac{\sqrt{1 + n(n+1)}}{|1 + r_n^{(1)}(\mathrm{i}sR)|^2} |sR|^2 |u_{1n}^m|^2 + \sum_{|m| \leq n}^{\text{even}} \frac{|1 + r_n^{(1)}(\mathrm{i}sR)|^2}{\sqrt{1 + n(n+1)}} \frac{|u_{2n}^m|^2}{|sR|^2} \right]^{1/2} \\ &\times \left[ \sum_{|m| \leq n}^{\text{odd}} \frac{1}{\sqrt{1 + n(n+1)}} |v_{1n}^m|^2 + \sum_{|m| \leq n}^{\text{even}} \sqrt{1 + n(n+1)} |v_{2n}^m|^2 \right]^{1/2}. \end{split}$$

By Lemma C.3, we have

$$\begin{split} \frac{\sqrt{1+n(n+1)}}{|1+r_n^{(1)}(\mathrm{i} s R)|^2} \, |sR|^2 \, |u_{1n}^m|^2 &= \frac{1}{\sqrt{1+n(n+1)}} \frac{1+n(n+1)}{|1+r_n^{(1)}(\mathrm{i} s R)|^2} |sR|^2 |u_{1n}^m|^2 \\ &\lesssim \frac{1}{\sqrt{1+n(n+1)}} \, |u_{1n}^m|^2, \end{split}$$

and

$$\begin{split} \frac{|1+r_n^{(1)}(\mathrm{i} sR)|^2}{\sqrt{1+n(n+1)}} \frac{|u_{2n}^m|^2}{|sR|^2} &= \sqrt{1+n(n+1)} \frac{|1+r_n^{(1)}(\mathrm{i} sR)|^2}{1+n(n+1)} \frac{|u_{2n}^m|^2}{|sR|^2} \\ &\lesssim \sqrt{1+n(n+1)} \, |u_{2n}^m|^2. \end{split}$$

Combining the above estimates and using the expressions of the norms in Appendix D yields

$$|\langle \mathscr{B}[oldsymbol{u}], oldsymbol{v}
angle| \lesssim \|oldsymbol{u}\|_{oldsymbol{H}^{-1/2}(\operatorname{curl},\,\Gamma_{oldsymbol{p}}^{+})} \|oldsymbol{v}\|_{oldsymbol{H}^{-1/2}(\operatorname{curl},\,\Gamma_{oldsymbol{p}}^{+})},$$

which completes the proof.  $\square$ 

Another important property of the EtM Calderon operator is stated as follows. LEMMA 2.2. It holds

(2.19) 
$$\operatorname{Re} \langle \mathscr{B}[\boldsymbol{u}], \boldsymbol{u} \rangle \geq 0, \quad \forall \, \boldsymbol{u} \in \boldsymbol{H}^{-1/2}(\operatorname{curl}, \Gamma_R^+).$$

*Proof.* From (2.18), we obtain

$$(2.20) -\langle \mathscr{B}[\boldsymbol{u}], \boldsymbol{u} \rangle = \sum_{|m| \le n}^{\text{odd}} \frac{sR}{1 + r_n^{(1)}(\text{i}sR)} |u_{1n}^m|^2 + \sum_{|m| \le n}^{\text{even}} \frac{1 + r_n^{(1)}(\text{i}sR)}{sR} |u_{2n}^m|^2.$$

By Lemma C.1 and Lemma C.2,

$$s_1 \text{Re}(1 + r_n^{(1)}(isR)) \le 0, \quad s_2 \text{Im}(1 + r_n^{(1)}(isR)) \le 0.$$

Taking the real part of (2.20) gives

$$-\operatorname{Re}\langle \mathscr{B}[\boldsymbol{u}], \boldsymbol{u} \rangle = \sum_{|m| \le n}^{\operatorname{odd}} \frac{R\left(s_1 \operatorname{Re}(1 + r_n^{(1)}(\mathrm{i}sR)) + s_2 \operatorname{Im}(1 + r_n^{(1)}(\mathrm{i}sR))\right)}{|1 + r_n^{(1)}(\mathrm{i}sR)|^2} |u_{1n}^m|^2 + \sum_{|m| \le n}^{\operatorname{even}} \frac{\left(s_1 \operatorname{Re}(1 + r_n^{(1)}(\mathrm{i}sR)) + s_2 \operatorname{Im}(1 + r_n^{(1)}(\mathrm{i}sR))\right)}{|s|^2 R} |u_{2n}^m|^2 \le 0,$$

which completes the proof.  $\Box$ 

With the aid of the frequency domain EtM Calderon operator, we obtain the following TBC imposed upon the hemisphere  $\Gamma_R^+$  in the s-domain:

(2.21) 
$$\mathscr{B}[\check{\mathbf{E}}_{\Gamma_{p}^{+}}] = \check{\mathbf{H}} \times \mathbf{e}_{\rho},$$

which maps the tangential component of the electric field to the tangential trace of the magnetic field. Taking the inverse Laplace transform of (2.21) yields the TBC in the time-domain:

$$(2.22) \hspace{1cm} \mathscr{T}[\pmb{E}_{\Gamma_R^+}] = \pmb{H} \times \pmb{e}_\rho, \quad \text{where} \ \ \mathscr{T} := \mathscr{L}^{-1} \circ \mathscr{B} \circ \mathscr{L}.$$

Equivalently, we may eliminate the magnetic field and obtain an alternative transparent boundary condition in the s-domain:

(2.23) 
$$s^{-1}(\nabla \times \boldsymbol{E}) \times \boldsymbol{n} + \mathscr{B}[\boldsymbol{E}_{\Gamma_{n}^{+}}] = 0 \quad \text{on } \Gamma_{R}^{+}.$$

Correspondingly, by taking the inverse Laplace transform of (2.23), we may derive an alternative TBC in the time domain:

$$(2.24) (\nabla \times \mathbf{E}) \times \mathbf{n} + \mathscr{C}[\mathbf{E}_{\Gamma_{\mathbf{p}}^{+}}] = 0 \text{on } \Gamma_{R}^{+}, \text{where } \mathscr{C} = \mathscr{L}^{-1} \circ s\mathscr{B} \circ \mathscr{L}.$$

- **3.** Analysis of two auxiliary problems. In this section, we make necessary preparations for the proof of the main results by considering two auxiliary problems pertinent to (2.5).
- **3.1. Time-harmonic Maxwell equations with a complex wavenumber.** This subsection is devoted to the mathematical study of a time-harmonic Maxwell scattering problem with a complex wavenumber, which may be viewed as a frequency version of the initial-boundary value problem of the Maxwell equations under the Laplace transform.

Consider the auxiliary boundary value problem:

(3.1) 
$$\begin{cases} \nabla \times ((s\mu)^{-1}\nabla \times \boldsymbol{u}) + s\varepsilon \boldsymbol{u} = \boldsymbol{j} & \text{in } \Omega, \\ \boldsymbol{n} \times \boldsymbol{u} = 0 & \text{on } S, \\ s^{-1}(\nabla \times \boldsymbol{u}) \times \boldsymbol{n} + \mathscr{B}[\boldsymbol{u}_{\Gamma}] = 0 & \text{on } \Gamma_{R}^{+}. \end{cases}$$

where  $s = s_1 + \mathrm{i} s_2$  with  $s_1, s_2 \in \mathbb{R}, s_1 > 0$ , and the applied current density  $\boldsymbol{j}$  is assumed to be supported in  $D \subset \Omega$ .

By multiplying a test function  $\mathbf{v} \in \mathbf{H}_S(\text{curl}, \Omega)$ , which is defined in Appendix D, and integrating by parts, we arrive at the variational formulation of (3.1): find  $\mathbf{u} \in \mathbf{H}_S(\text{curl}, \Omega)$  such that

(3.2) 
$$a_{\mathrm{TH}}(\boldsymbol{u},\boldsymbol{v}) = \int_{\Omega} \boldsymbol{j} \cdot \bar{\boldsymbol{v}} \, \mathrm{d}\boldsymbol{r} \quad \text{for all } \boldsymbol{v} \in \boldsymbol{H}_{S}(\text{curl},\Omega),$$

where the sesquilinear form

$$(3.3) \quad a_{\mathrm{TH}}(\boldsymbol{u}, \boldsymbol{v}) = \int_{\Omega} (s\mu)^{-1} (\nabla \times \boldsymbol{u}) \cdot (\nabla \times \bar{\boldsymbol{v}}) \, \mathrm{d}\boldsymbol{r} + \int_{\Omega} s\varepsilon \, \boldsymbol{u} \cdot \bar{\boldsymbol{v}} \, \mathrm{d}\boldsymbol{r} + \langle \mathscr{B}[\boldsymbol{u}_{\Gamma_{R}^{+}}], \boldsymbol{v}_{\Gamma_{R}^{+}} \rangle.$$

THEOREM 3.1. The variational problem (3.2) has a unique solution  $\mathbf{u} \in \mathbf{H}_S(\text{curl}, \Omega)$ , such that

$$\|\nabla\times\boldsymbol{u}\|_{\boldsymbol{L}^{2}(\Omega)}+\|s\boldsymbol{u}\|_{\boldsymbol{L}^{2}(\Omega)}\lesssim s_{1}^{-1}\|s\boldsymbol{j}\|_{\boldsymbol{L}^{2}(\Omega)},$$

where  $s = s_1 + is_2$  with  $s_1, s_2 \in \mathbb{R}$  and  $s_1 > 0$ .

*Proof.* It suffices to show the coercivity of  $a_{\rm TH}$ , since its continuity follows directly from the Cauchy-Schwarz inequality, Lemma 2.1, and Lemma D.1. A simple calculation yields

(3.4) 
$$a_{\mathrm{TH}}(\boldsymbol{u}, \boldsymbol{u}) = \int_{\Omega} (s\mu)^{-1} |\nabla \times \boldsymbol{u}|^2 \, \mathrm{d}\boldsymbol{r} + \int_{\Omega} s\varepsilon |\boldsymbol{u}|^2 \, \mathrm{d}\boldsymbol{r} + \langle \mathscr{B}[\boldsymbol{u}_{\Gamma_R^+}], \boldsymbol{u}_{\Gamma_R^+} \rangle.$$

Taking the real part of (3.4) and using Lemma 2.2, we get

(3.5) 
$$\operatorname{Re}\left\{a_{\mathrm{TH}}(\boldsymbol{u},\boldsymbol{u})\right\} \gtrsim \frac{s_1}{|s|^2} \left(\|\nabla \times \boldsymbol{u}\|_{\boldsymbol{L}^2(\Omega)}^2 + \|s\boldsymbol{u}\|_{\boldsymbol{L}^2(\Omega)}^2\right).$$

It follows from the Lax-Milgram lemma that the variational problem (3.2) has a unique solution  $u \in H_S(\text{curl}, \Omega)$ . Moreover, we have from (3.2) that

(3.6) 
$$|a_{\mathrm{TH}}(\boldsymbol{u}, \boldsymbol{u})| \leq |s|^{-1} ||\boldsymbol{j}||_{L^{2}(\Omega)} ||s\boldsymbol{u}||_{L^{2}(\Omega)}.$$

Combining (3.5)-(3.6) leads to

$$\|\nabla \times \boldsymbol{u}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|s\boldsymbol{u}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \lesssim s_{1}^{-1} \|s\boldsymbol{j}\|_{\boldsymbol{L}^{2}(\Omega)} \|s\boldsymbol{u}\|_{\boldsymbol{L}^{2}(\Omega)},$$

which completes the proof after applying the Cauchy-Schwarz inequality.  $\square$ 

It is noteworthy that Theorem 3.1 gives the stability estimate with an explicit dependence on the complex wavenumber s corresponding to a lossy medium with  $s_1 > 0$ . However, it is challenging to obtain a similar estimate for a lossless medium with  $s_1 = 0$ .

**3.2.** A second auxiliary problem. Before considering the reduced problem (2.5), we need to study an auxiliary initial-boundary value problem and establish its well-posedness and stability. Consider the system of time-domain Maxwell equations:

(3.7) 
$$\begin{cases} \nabla \times \boldsymbol{U} + \mu \partial_t \boldsymbol{V} = 0, & \nabla \times \boldsymbol{V} - \varepsilon \partial_t \boldsymbol{U} = 0 & \text{in } \Omega, \ t > 0, \\ \boldsymbol{n} \times \boldsymbol{U} = 0 & \text{on } S \cup \Gamma_R^+, \ t > 0, \\ \boldsymbol{U}|_{t=0} = \boldsymbol{E}_0, \quad \boldsymbol{V}|_{t=0} = \boldsymbol{H}_0 & \text{in } \Omega, \end{cases}$$

where  $E_0, H_0$  are assumed to be compactly supported in D as before.

Let  $\check{\boldsymbol{U}} = \mathcal{L}(\boldsymbol{U})$  and  $\check{\boldsymbol{V}} = \mathcal{L}(\boldsymbol{V})$ . Taking the Laplace transform of (3.7), we obtain the boundary value problem:

(3.8) 
$$\begin{cases} \nabla \times \left( (s\mu)^{-1} \nabla \times \boldsymbol{\check{U}} \right) + s\varepsilon \, \boldsymbol{\check{U}} = \boldsymbol{\check{j}} & \text{in } \Omega, \\ \boldsymbol{n} \times \boldsymbol{\check{U}} = 0 & \text{on } S \cup \Gamma_R^+, \end{cases}$$

where the current density function

$$\ddot{\boldsymbol{j}} = \varepsilon \boldsymbol{E}_0 + s^{-1} \nabla \times \boldsymbol{H}_0.$$

The variational formulation for (3.8) is to find  $\check{\boldsymbol{U}} \in \boldsymbol{H}_0(\operatorname{curl},\Omega)$  such that

(3.9) 
$$a_{\mathrm{AP}}(\check{\boldsymbol{U}},\boldsymbol{v}) = \int_{\Omega} \check{\boldsymbol{j}} \cdot \bar{\boldsymbol{v}} \, \mathrm{d}\boldsymbol{r} \quad \text{for all } \boldsymbol{v} \in \boldsymbol{H}_{0}(\mathrm{curl},\Omega),$$

where the sesquilinear form

(3.10) 
$$a_{\rm AP}(\boldsymbol{u},\boldsymbol{v}) = \int_{\Omega} (s\mu)^{-1} (\nabla \times \boldsymbol{u}) \cdot (\nabla \times \bar{\boldsymbol{v}}) \, \mathrm{d}\boldsymbol{r} + \int_{\Omega} s\varepsilon \boldsymbol{u} \cdot \bar{\boldsymbol{v}} \, \mathrm{d}\boldsymbol{r}.$$

Following the same proof as in Theorem 3.1, we can show the well-posedness of the variational problem (3.9) and its stability, as stated below.

LEMMA 3.2. The variational problem (3.9) has a unique solution  $\check{\boldsymbol{U}} \in \boldsymbol{H}_0(\operatorname{curl},\Omega)$ . Moreover, we have the estimate

$$\|\nabla \times \check{\boldsymbol{U}}\|_{\boldsymbol{L}^{2}(\Omega)} + \|s\check{\boldsymbol{U}}\|_{\boldsymbol{L}^{2}(\Omega)} \lesssim s_{1}^{-1}|s|\|\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)} + s_{1}^{-1}\|\nabla \times \boldsymbol{H}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}.$$

Theorem 3.3. The auxiliary problem (3.7) has a unique solution (U, V), which satisfies the stability estimates

$$\begin{split} & \|\boldsymbol{U}\|_{\boldsymbol{L}^2(\Omega)} + \|\boldsymbol{V}\|_{\boldsymbol{L}^2(\Omega)} \lesssim \|\boldsymbol{E}_0\|_{\boldsymbol{L}^2(\Omega)} + \|\boldsymbol{H}_0\|_{\boldsymbol{L}^2(\Omega)}, \\ & \|\partial_t \boldsymbol{U}\|_{\boldsymbol{L}^2(\Omega)} + \|\partial_t \boldsymbol{V}\|_{\boldsymbol{L}^2(\Omega)} \lesssim \|\nabla \times \boldsymbol{E}_0\|_{\boldsymbol{L}^2(\Omega)} + \|\nabla \times \boldsymbol{H}_0\|_{\boldsymbol{L}^2(\Omega)}, \\ & \|\partial_t^2 \boldsymbol{U}\|_{\boldsymbol{L}^2(\Omega)} + \|\partial_t^2 \boldsymbol{V}\|_{\boldsymbol{L}^2(\Omega)} \lesssim \|\nabla \times (\nabla \times \boldsymbol{E}_0)\|_{\boldsymbol{L}^2(\Omega)} + \|\nabla \times (\nabla \times \boldsymbol{H}_0)\|_{\boldsymbol{L}^2(\Omega)}. \end{split}$$

*Proof.* Let  $\check{\boldsymbol{U}}=\mathscr{L}(\boldsymbol{U})$  and  $\check{\boldsymbol{V}}=\mathscr{L}(\boldsymbol{V})$  as before. Taking the Laplace transform of (3.7) leads to

(3.11) 
$$\begin{cases} \nabla \times \boldsymbol{\breve{U}} + s\mu \, \boldsymbol{\breve{V}} = \mu \, \boldsymbol{H}_0, \ \nabla \times \boldsymbol{\breve{V}} - s\varepsilon \, \boldsymbol{\breve{U}} = -\varepsilon \, \boldsymbol{E}_0 & \text{in } \Omega, \\ \boldsymbol{n} \times \boldsymbol{\breve{U}} = 0 & \text{on } S \cap \Gamma_B^+. \end{cases}$$

It follows from Lemma 3.2 that

$$\|\nabla \times \check{\boldsymbol{U}}\|_{\boldsymbol{L}^{2}(\Omega)} + \|s\check{\boldsymbol{U}}\|_{\boldsymbol{L}^{2}(\Omega)} \lesssim s_{1}^{-1}|s|\|\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)} + s_{1}^{-1}\|\nabla \times \boldsymbol{H}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}.$$

By (3.11), we have

$$\|\nabla \times \breve{\boldsymbol{V}}\|_{\boldsymbol{L}^{2}(\Omega)} + \|s\breve{\boldsymbol{V}}\|_{\boldsymbol{L}^{2}(\Omega)} \lesssim (1 + s_{1}^{-1}|s|) \|\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)} + \|\boldsymbol{H}_{0}\|_{\boldsymbol{L}^{2}(\Omega)} + s_{1}^{-1}\|\nabla \times \boldsymbol{H}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}.$$

It follows from Lemma A.2 that  $\check{\boldsymbol{U}}$  and  $\check{\boldsymbol{V}}$  are holomorphic functions of s, and the inverse Laplace transform of  $\check{\boldsymbol{U}}$  and  $\check{\boldsymbol{V}}$  exist and are supported in  $[0,\infty)$ .

We next prove the stability. Define the energy function

$$e_1(t) = \|\varepsilon^{1/2} \boldsymbol{U}(\cdot, t)\|_{\boldsymbol{L}^2(\Omega)}^2 + \|\mu^{1/2} \boldsymbol{V}(\cdot, t)\|_{\boldsymbol{L}^2(\Omega)}^2.$$

It follows from (3.7) and integration by parts that

$$e_{1}(t) - e_{1}(0) = \int_{0}^{t} e'_{1}(\tau) d\tau = 2\operatorname{Re} \int_{0}^{t} \int_{\Omega} \left( \varepsilon \partial_{t} \boldsymbol{U} \cdot \bar{\boldsymbol{U}} + \mu \partial_{t} \boldsymbol{V} \cdot \bar{\boldsymbol{V}} \right) d\boldsymbol{r} d\tau$$

$$= 2\operatorname{Re} \int_{0}^{t} \int_{\Omega} \left( (\nabla \times \boldsymbol{V}) \cdot \bar{\boldsymbol{U}} - (\nabla \times \boldsymbol{U}) \cdot \bar{\boldsymbol{V}} \right) d\boldsymbol{r} d\tau$$

$$= 2\operatorname{Re} \int_{0}^{t} \int_{\Omega} \left( (\nabla \times \bar{\boldsymbol{U}}) \cdot \boldsymbol{V} - (\nabla \times \boldsymbol{U}) \cdot \bar{\boldsymbol{V}} \right) d\boldsymbol{r} d\tau = 0.$$

Hence we have

$$\|\varepsilon^{1/2}\boldsymbol{U}(\cdot,t)\|_{\boldsymbol{L}^{2}(\Omega)}^{2}+\|\mu^{1/2}\boldsymbol{V}(\cdot,t)\|_{\boldsymbol{L}^{2}(\Omega)}^{2}=\|\varepsilon^{1/2}\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}^{2}+\|\mu^{1/2}\boldsymbol{H}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}^{2},$$

which implies

$$\|\boldsymbol{U}\|_{\boldsymbol{L}^{2}(\Omega)} + \|\boldsymbol{V}\|_{\boldsymbol{L}^{2}(\Omega)} \lesssim \|\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)} + \|\boldsymbol{H}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}.$$

Taking the first and second partial derivatives of (3.7) with respect to t yields

$$\begin{cases} \nabla \times \partial_t \boldsymbol{U} + \mu \partial_t^2 \boldsymbol{V} = \boldsymbol{0}, \ \nabla \times \partial_t \boldsymbol{V} - \varepsilon \partial_t^2 \boldsymbol{U} = 0 & \text{in } \Omega, \ t > 0, \\ \boldsymbol{n} \times \partial_t \boldsymbol{U} = \boldsymbol{0} & \text{on } S \cup \Gamma_R^+, \ t > 0, \\ \partial_t \boldsymbol{U}|_{t=0} = \varepsilon^{-1} (\nabla \times \boldsymbol{H}_0), \ \partial_t \boldsymbol{V}|_{t=0} = -\mu^{-1} (\nabla \times \boldsymbol{E}_0) & \text{in } \Omega, \end{cases}$$

and

$$\begin{cases} \nabla \times \partial_t^2 \boldsymbol{U} + \mu \partial_t^3 \boldsymbol{V} = 0, \ \nabla \times \partial_t^2 \boldsymbol{V} - \varepsilon \partial_t^3 \boldsymbol{U} = 0 & \text{in } \Omega, \ t > 0, \\ \boldsymbol{n} \times \partial_t^2 \boldsymbol{U} = 0 & \text{on } S \cup \Gamma_R^+, \ t > 0, \\ \partial_t^2 \boldsymbol{U}|_{t=0} = -(\varepsilon \mu)^{-1} (\nabla \times (\nabla \times \boldsymbol{E}_0)) & \text{in } \Omega, \\ \partial_t^2 \boldsymbol{V}|_{t=0} = -(\varepsilon \mu)^{-1} (\nabla \times (\nabla \times \boldsymbol{H}_0)) & \text{in } \Omega. \end{cases}$$
Hering the energy functions

Considering the energy functions

$$e_2(t) = \|\varepsilon^{1/2} \partial_t \boldsymbol{U}(\cdot, t)\|_{\boldsymbol{L}^2(\Omega)}^2 + \|\mu^{1/2} \partial_t \boldsymbol{V}(\cdot, t)\|_{\boldsymbol{L}^2(\Omega)}^2,$$

and

$$e_3(t) = \|\varepsilon^{1/2}\partial_t^2 \mathbf{U}(\cdot, t)\|_{\mathbf{L}^2(\Omega)}^2 + \|\mu^{1/2}\partial_t^2 \mathbf{V}(\cdot, t)\|_{\mathbf{L}^2(\Omega)}^2$$

for the above two problems, respectively, we can follow the same steps for proving the first inequality to derive the other two inequalities.  $\square$ 

- **4.** The reduced Maxwell equations. In this section, we derive the main results of this work, which include the well-posedness of the reduced problem (2.5) and the related *a priori* estimates.
- **4.1. Well-posedness.** Let e = E U and h = H V. It follows from (2.5) and (3.7) that e and h satisfy the system:

(4.1) 
$$\begin{cases} \nabla \times \boldsymbol{e} + \mu \partial_t \boldsymbol{h} = 0, \ \nabla \times \boldsymbol{h} - \varepsilon \partial_t \boldsymbol{e} = \boldsymbol{J} & \text{in } \Omega, \ t > 0, \\ \boldsymbol{n} \times \boldsymbol{e} = 0 & \text{on } S, \ t > 0, \\ \boldsymbol{e}|_{t=0} = 0, \ \boldsymbol{h}|_{t=0} = 0 & \text{in } \Omega, \\ \mathscr{T} \boldsymbol{e}_{\Gamma_R^+} = \boldsymbol{h} \times \boldsymbol{n} + \boldsymbol{V} \times \boldsymbol{n} & \text{on } \Gamma_R^+, \ t > 0. \end{cases}$$

Let  $\check{e} = \mathcal{L}(e)$  and  $\check{h} = \mathcal{L}(h)$ . Taking the Laplace transform of (4.1) and eliminating  $\check{h}$ , we obtain

(4.2) 
$$\begin{cases} \nabla \times ((s\mu)^{-1})\nabla \times \check{\boldsymbol{e}}) + s\varepsilon \check{\boldsymbol{e}} = -\check{\boldsymbol{J}} & \text{in } \Omega, \\ \boldsymbol{n} \times \check{\boldsymbol{e}} = 0 & \text{on } S, \\ s^{-1}(\nabla \times \check{\boldsymbol{e}}) \times \boldsymbol{n} + \mathscr{B}\check{\boldsymbol{e}}_{\Gamma_{R}^{+}} = \check{\boldsymbol{V}} \times \boldsymbol{n} & \text{on } \Gamma_{R}^{+}. \end{cases}$$

Lemma 4.1. The problem (4.2) has a unique weak solution  $\check{e} \in H_S(\text{curl}, \Omega)$ , which satisfies

$$\begin{split} \|\nabla \times \check{\boldsymbol{e}}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|s\check{\boldsymbol{e}}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \lesssim s_{1}^{-1} \Big( \|s\check{\boldsymbol{J}}\|_{\boldsymbol{L}^{2}(\Omega)} + \|s\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_{R}^{+})} \\ + \||s|^{2}\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_{R}^{+})} \Big). \end{split}$$

*Proof.* The well-posedness of the solution  $\check{e} \in H_S(\text{curl}, \Omega)$  follows directly from Theorem 3.1. Moreover, we have

$$a_{\mathrm{TH}}(reve{e},reve{e}) = \langle reve{V} imes m{n}, reve{e}_{\Gamma_R^+} 
angle - \int_{\Omega} reve{J} \cdot ar{reve{e}} \, \mathrm{d}m{r}.$$

By the coercivity of  $a_{TH}$  in (3.5) and the trace theorem in Lemma D.1,

$$\begin{split} \frac{s_1}{|s|^2} \Big( \|\nabla \times \check{\boldsymbol{e}}\|_{\boldsymbol{L}^2(\Omega)}^2 + \|s\check{\boldsymbol{e}}\|_{\boldsymbol{L}^2(\Omega)}^2 \Big) &\lesssim \|s^{-1} \check{\boldsymbol{J}}\|_{\boldsymbol{L}^2(\Omega)} \|s\check{\boldsymbol{e}}\|_{\boldsymbol{L}^2(\Omega)} \\ &+ \|\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\,\Gamma_R^+)} \|\check{\boldsymbol{e}}\|_{\boldsymbol{H}(\operatorname{curl},\,\Omega)} \\ &\lesssim \|s^{-1} \check{\boldsymbol{J}}\|_{\boldsymbol{L}^2(\Omega)} \|s\check{\boldsymbol{e}}\|_{\boldsymbol{L}^2(\Omega)} + \|\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\,\Gamma_R^+)} \|\nabla \times \check{\boldsymbol{e}}\|_{\boldsymbol{L}^2(\Omega)} \\ &+ \|s^{-1} \check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\,\Gamma_R^+)} \|s\check{\boldsymbol{e}}\|_{\boldsymbol{L}^2(\Omega)}, \end{split}$$

which completes the proof.  $\square$ 

To show the well-posedness of the reduced problem (2.5), we make the following assumptions for the initial and boundary data:

(4.3) 
$$E_0, H_0 \in H(\text{curl}, \Omega), J \in H^1(0, T; L^2(\Omega)), J|_{t=0} = 0.$$

Theorem 4.2. The problem has a unique solution (E, H) such that

$$E \in L^2(0,T; \boldsymbol{H}_S(\text{curl}, \Omega)) \cap \boldsymbol{H}^1(0,T; \boldsymbol{L}^2(\Omega)),$$
  
 $\boldsymbol{H} \in L^2(0,T; \boldsymbol{H}(\text{curl}, \Omega)) \cap \boldsymbol{H}^1(0,T; \boldsymbol{L}^2(\Omega)).$ 

Moreover, they satisfy

$$\max_{t \in [0,T]} \left( \|\partial_t \boldsymbol{E}\|_{\boldsymbol{L}^2(\Omega)} + \|\nabla \times \boldsymbol{E}\|_{\boldsymbol{L}^2(\Omega)} + \|\partial_t \boldsymbol{H}\|_{\boldsymbol{L}^2(\Omega)} + \|\nabla \times \boldsymbol{H}\|_{\boldsymbol{L}^2(\Omega)} \right)$$

$$\lesssim \|\boldsymbol{E}_0\|_{\boldsymbol{H}(\operatorname{curl},\Omega)} + \|\boldsymbol{H}_0\|_{\boldsymbol{H}(\operatorname{curl},\Omega)} + \|\boldsymbol{J}\|_{\boldsymbol{H}^1(0,T;\boldsymbol{L}^2(\Omega))}.$$

*Proof.* Recall the decomposition E = U + e and H = V + h where (U, V) satisfy (3.7) and (e, h) satisfy (4.1). Since

$$\int_{0}^{T} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}\boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) dt$$

$$\leq \int_{0}^{T} e^{-2s_{1}(t-T)} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}\boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) dt$$

$$= e^{2s_{1}T} \int_{0}^{T} e^{-2s_{1}t} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}\boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) dt$$

$$\lesssim \int_{0}^{\infty} e^{-2s_{1}t} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}\boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) dt,$$

it suffices to estimate the integral

$$\int_0^\infty e^{-2s_1t} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^2(\Omega)}^2 + \|\partial_t \boldsymbol{e}\|_{\boldsymbol{L}^2(\Omega)}^2 \right) dt.$$

Taking the Laplace transform of (4.1) yields

$$\begin{cases} \nabla \times \check{\boldsymbol{e}} + s\mu\check{\boldsymbol{h}} = 0, \quad \nabla \times \check{\boldsymbol{h}} - s\varepsilon\check{\boldsymbol{e}} = \check{\boldsymbol{J}} & \text{in } \Omega, \\ \boldsymbol{n} \times \check{\boldsymbol{e}} = 0 & \text{on } S, \\ \mathscr{B}\check{\boldsymbol{e}}_{\Gamma_R^+} = \check{\boldsymbol{h}} \times \boldsymbol{n} + \check{\boldsymbol{V}} \times \boldsymbol{n} & \text{on } \Gamma_R^+ \end{cases}$$

By Lemma 4.1,

$$\|\nabla \times \check{\boldsymbol{e}}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|s\check{\boldsymbol{e}}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \lesssim s_{1}^{-1} \Big(\|s\check{\boldsymbol{J}}\|_{\boldsymbol{L}^{2}(\Omega)} + \|s\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_{R}^{+})} + \||s|^{2}\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_{R}^{+})}\Big).$$

$$(4.5)$$

By (4.4),

$$\|\nabla \times \check{\boldsymbol{h}}\|_{L^{2}(\Omega)}^{2} + \|s\check{\boldsymbol{h}}\|_{L^{2}(\Omega)}^{2} \lesssim s_{1}^{-1} \Big( \|\check{\boldsymbol{J}}\|_{L^{2}(\Omega)} + \|s\check{\boldsymbol{J}}\|_{L^{2}(\Omega)} + \|s\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_{R}^{+})} + \||s|^{2}\check{\boldsymbol{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_{R}^{+})} \Big).$$

$$(4.6)$$

It follows from [33, Lemma 44.1] that  $(\check{e}, \check{h})$  are holomorphic functions of s on the half plane  $s_1 > \gamma > 0$ , where  $\gamma$  is any positive constant. Hence we have from Lemma A.2 that the inverse Laplace transform of  $\check{e}$  and  $\check{h}$  exist and are supported in  $[0, \infty]$ .

Denote by  $e = \mathcal{L}^{-1}(\check{e})$  and  $h = \mathcal{L}^{-1}(\check{h})$ . Since  $\check{e} = \mathcal{L}(e) = \mathcal{F}(e^{-s_1t}e)$ , where  $\mathscr{F}$  is the Fourier transform in  $s_2$ , we have from the Parseval identity and (4.5) that

$$\int_{0}^{\infty} e^{-2s_{1}t} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}\boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) dt = 2\pi \int_{-\infty}^{\infty} \left( \|\nabla \times \boldsymbol{\check{e}}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|s\boldsymbol{\check{e}}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) ds_{2}$$

$$\lesssim s_{1}^{-2} \int_{-\infty}^{\infty} \|s\boldsymbol{\check{J}}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} ds_{2} + s_{1}^{-2} \int_{-\infty}^{\infty} \left( \|s\boldsymbol{\check{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\Gamma_{R}^{+})}^{2} \right) ds_{2}.$$

$$+ \||s|^{2} \boldsymbol{\check{V}} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\Gamma_{R}^{+})}^{2} \right) ds_{2}.$$

Since  $J|_{t=0} = 0$  in  $\Omega$ ,  $V \times n|_{t=0} = \partial_t (V \times n)|_{t=0} = 0$  on  $\Gamma_R^+$ , we have  $\mathcal{L}(\partial_t J) = s \check{J}$  in  $\Omega$  and  $\mathcal{L}(\partial_t (V \times n)) = s \check{V} \times n$  on  $\Gamma_R^+$ . It is easy to note that

$$|s|^2 \breve{\boldsymbol{V}} \times \boldsymbol{n} = (2s_1 - s)s \breve{\boldsymbol{V}} \times \boldsymbol{n} = 2s_1 \mathcal{L}(\partial_t (\boldsymbol{V} \times \boldsymbol{n})) - \mathcal{L}(\partial_t^2 (\boldsymbol{V} \times \boldsymbol{n}))$$
 on  $\Gamma_R^+$ .

Hence we have

$$\int_{0}^{\infty} e^{-2s_{1}t} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}\boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) dt$$

$$\lesssim s_{1}^{-2} \int_{-\infty}^{\infty} \left( \|\mathcal{L}(\partial_{t}\boldsymbol{J})\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\mathcal{L}(\partial_{t}^{2}\boldsymbol{V} \times \boldsymbol{n})\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\Gamma_{R}^{+})}^{2} \right) ds_{2}$$

$$+ (1 + s_{1}^{-2}) \int_{-\infty}^{\infty} \|\mathcal{L}(\partial_{t}\boldsymbol{V} \times \boldsymbol{n})\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\Gamma_{R}^{+})}^{2} ds_{2}.$$

Using the Parseval identity again gives

$$\int_{0}^{\infty} e^{-2s_{1}t} \left( \|\nabla \times \boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}\boldsymbol{e}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) dt$$

$$\lesssim s_{1}^{-2} \int_{0}^{\infty} e^{-2s_{1}t} \left( \|\partial_{t}\boldsymbol{J}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\partial_{t}^{2}\boldsymbol{V} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\Gamma_{R}^{+})}^{2} \right) dt$$

$$+ (1 + s_{1}^{-2}) \int_{0}^{\infty} e^{-2s_{1}t} \|\partial_{t}\boldsymbol{V} \times \boldsymbol{n}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div},\Gamma_{R}^{+})}^{2} dt,$$

which shows that

$$e \in L^2(0,T; \boldsymbol{H}_S(\text{curl},\Omega)) \cap \boldsymbol{H}^1(0,T; \boldsymbol{L}^2(\Omega)).$$

Similarly, we can show from (4.6) that

$$\boldsymbol{h} \in \boldsymbol{L}^2(0,T;\boldsymbol{H}(\operatorname{curl},\Omega)) \cap \boldsymbol{H}^1(0,T;\boldsymbol{L}^2(\Omega)).$$

We next prove the stability. Let  $\tilde{\boldsymbol{E}}$  be the extension of  $\boldsymbol{E}$  with respect to t in  $\mathbb{R}$  such that  $\tilde{\boldsymbol{E}} = 0$  outside the interval [0, t]. By Lemma A.1 and Lemma 2.2, we get

$$\operatorname{Re} \int_{0}^{t} e^{-2s_{1}t} \int_{\Gamma_{R}^{+}} \mathscr{T}[\boldsymbol{E}_{\Gamma_{R}^{+}}] \cdot \bar{\boldsymbol{E}}_{\Gamma_{R}^{+}} \, \mathrm{d}\gamma \, \mathrm{d}t = \operatorname{Re} \int_{\Gamma_{R}^{+}} \int_{0}^{\infty} e^{-2s_{1}t} \mathscr{T}[\tilde{\boldsymbol{E}}_{\Gamma_{R}^{+}}] \cdot \bar{\tilde{\boldsymbol{E}}}_{\Gamma_{R}^{+}} \, \mathrm{d}t \, \mathrm{d}\gamma$$
$$= \frac{1}{2\pi} \int_{-\infty}^{\infty} \operatorname{Re} \langle \mathscr{B}[\check{\boldsymbol{E}}_{\Gamma_{R}^{+}}](s), \, \check{\boldsymbol{E}}_{\Gamma_{R}^{+}}(s) \rangle \, \mathrm{d}s_{2} \geq 0,$$

which yields after taking  $s_1 \to 0$  that

(4.7) 
$$\operatorname{Re} \int_0^t \int_{\Gamma_R^+} \mathscr{T}[\boldsymbol{E}_{\Gamma_R^+}] \cdot \bar{\boldsymbol{E}}_{\Gamma_R^+} \, \mathrm{d}\gamma \mathrm{d}t \ge 0.$$

For any 0 < t < T, consider an energy function

$$e(t) = \|\varepsilon^{1/2}\boldsymbol{E}(\cdot,t)\|_{\boldsymbol{L}^2(\Omega)}^2 + \|\mu^{1/2}\boldsymbol{H}(\cdot,t)\|_{\boldsymbol{L}^2(\Omega)}^2.$$

It is easy to note that

(4.8) 
$$\int_{0}^{t} u'(t)dt = \left( \|\varepsilon^{1/2} \boldsymbol{E}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\mu^{1/2} \boldsymbol{H}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right) - \left( \|\varepsilon^{1/2} \boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \|\mu^{1/2} \boldsymbol{H}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} \right).$$

On the other hand, we have from (2.5) and (4.7) that

$$\int_{0}^{t} e'(t) dt = 2 \operatorname{Re} \int_{0}^{t} \int_{\Omega} \left( \varepsilon \partial_{t} \boldsymbol{E} \cdot \bar{\boldsymbol{E}} + \mu \partial_{t} \boldsymbol{H} \cdot \bar{\boldsymbol{H}} \right) d\boldsymbol{r} dt 
= 2 \operatorname{Re} \int_{0}^{t} \int_{\Omega} \left( (\nabla \times \boldsymbol{H}) \cdot \bar{\boldsymbol{E}} - (\nabla \times \boldsymbol{E}) \cdot \bar{\boldsymbol{H}} \right) d\boldsymbol{r} dt - 2 \operatorname{Re} \int_{0}^{t} \int_{\Omega} \boldsymbol{J} \cdot \bar{\boldsymbol{E}} d\boldsymbol{r} dt 
= -2 \operatorname{Re} \int_{0}^{t} \int_{\Gamma} \mathscr{T} [\boldsymbol{E}_{\Gamma_{R}^{+}}] \cdot \bar{\boldsymbol{E}}_{\Gamma_{R}^{+}} d\gamma dt - 2 \operatorname{Re} \int_{0}^{t} \int_{\Omega} \boldsymbol{J} \cdot \bar{\boldsymbol{E}} d\boldsymbol{r} dt 
= -2 \operatorname{Re} \int_{0}^{t} \int_{\Gamma} \mathscr{T} [\boldsymbol{E}_{\Gamma_{R}^{+}}] \cdot \bar{\boldsymbol{E}}_{\Gamma_{R}^{+}} d\gamma dt - 2 \operatorname{Re} \int_{0}^{t} \int_{\Omega} \boldsymbol{J} \cdot \bar{\boldsymbol{E}} d\boldsymbol{r} dt 
\leq -2 \operatorname{Re} \int_{0}^{t} \int_{\Omega} \boldsymbol{J} \cdot \bar{\boldsymbol{E}} d\boldsymbol{r} dt \leq 2 \max_{t \in [0,T]} \|\boldsymbol{E}(\cdot,t)\|_{\boldsymbol{L}^{2}(\Omega)} \|\boldsymbol{J}\|_{\boldsymbol{L}^{1}(0,T;\boldsymbol{L}^{2}(\Omega))}.$$

Taking the derivative of (2.5) with repsect to t, we know that  $(\partial_t \mathbf{E}, \partial_t \mathbf{H})$  satisfy the same set of equations with the source and the initial condition replaced by  $\partial_t \mathbf{J}$ ,  $\partial_t \mathbf{E}|_{t=0} = \varepsilon^{-1} \nabla \times \mathbf{E}_0$ ,  $\partial_t \mathbf{H}|_{t=0} = -\mu^{-1} \nabla \times \mathbf{H}_0$ . Hence we may follow the same steps as above to obtain (4.9) for  $(\partial_t \mathbf{E}, \partial_t \mathbf{H})$ , which completes the proof after combining the above estimates.  $\square$ 

**4.2.** An *a priori* estimate. In what follows, we shall take a different route and study the Maxwell equations directly in the time domain. The goal is to derive *a priori* stability estimate for the electric field with a minimum regularity requirement for the data and an explicit dependence on the time variable.

Eliminating the magnetic field in (2.1) and using the transparent boundary condition (2.24), we consider the initial-boundary value problem in a bounded domain:

$$\begin{cases} \varepsilon \partial_t^2 \boldsymbol{E} = -\nabla \times (\mu^{-1} \nabla \times \boldsymbol{E}) - \boldsymbol{F} & \text{in } \Omega, \ t > 0, \\ \boldsymbol{E}|_{t=0} = \boldsymbol{E}_0, \quad \partial_t \boldsymbol{E}|_{t=0} = \boldsymbol{E}_1 & \text{in } \Omega, \\ \boldsymbol{n} \times \boldsymbol{E} = 0 & \text{on } S, \ t > 0, \\ (\nabla \times \boldsymbol{E}) \times \boldsymbol{n} + \mathscr{C}[\boldsymbol{E}_{\Gamma_R^+}] = 0 & \text{on } \Gamma_R^+, \ t > 0, \end{cases}$$

where

$$F = \partial_t J$$
,  $E_1 = \varepsilon^{-1} (\nabla \times H_0 - J_0)$ .

The variational problem is to find  $E \in H_S(\text{curl}, \Omega)$  for all t > 0 such that

$$\int_{\Omega} \varepsilon \partial_{t}^{2} \boldsymbol{E} \cdot \bar{\boldsymbol{w}} \, d\boldsymbol{r} = -\int_{\Omega} \mu^{-1} (\nabla \times \boldsymbol{E}) \cdot (\nabla \times \bar{\boldsymbol{w}}) \, d\boldsymbol{r}$$

$$-\langle \mathscr{C}[\boldsymbol{E}_{\Gamma_{R}^{+}}], \boldsymbol{w}_{\Gamma_{R}^{+}} \rangle - \int_{\Omega} \boldsymbol{F} \cdot \bar{\boldsymbol{w}} \, d\boldsymbol{r} \quad \text{for all } \boldsymbol{w} \in \boldsymbol{H}_{S}(\text{curl}, \Omega).$$

To show the stability of its solution, we follow the argument in [33] but with a careful study of the TBC. The following lemma is useful for the subsequent analysis. Lemma 4.3. Given  $\xi \geq 0$  and any  $\mathbf{E} \in L^2(0, \xi; \mathbf{H}^{-1/2}(\text{curl}, \Gamma))$ , there holds

$$\operatorname{Re} \int_0^\xi \int_{\Gamma_R^+} \left( \int_0^t \mathscr{C}[\boldsymbol{E}_{\Gamma_R^+}](\tau) \mathrm{d}\tau \right) \cdot \bar{\boldsymbol{E}}_{\Gamma_R^+}(t) \, \mathrm{d}\gamma \mathrm{d}t \geq 0.$$

*Proof.* Let  $\tilde{\boldsymbol{E}}$  be the extension of  $\boldsymbol{E}$  with respect to t in  $\mathbb{R}$  such that  $\tilde{\boldsymbol{E}} = 0$  outside the interval  $[0, \xi]$ . We obtain from (A.1), Lemma A.2, and Lemma 2.2 that

$$\begin{split} \int_{\Gamma_R^+} \int_0^{\xi} e^{-2s_1 t} \left( \int_0^t \mathscr{C}[\boldsymbol{E}_{\Gamma_R^+}](\tau) \mathrm{d}\tau \right) \cdot \bar{\boldsymbol{E}}_{\Gamma_R^+}(t) \, \mathrm{d}t \mathrm{d}\gamma \\ &= \int_{\Gamma_R^+} \int_0^{\infty} e^{-2s_1 t} \left( \int_0^t \mathscr{C}[\tilde{\boldsymbol{E}}_{\Gamma_R^+}](\tau) \mathrm{d}\tau \right) \cdot \bar{\tilde{\boldsymbol{E}}}_{\Gamma_R^+}(t) \, \mathrm{d}t \mathrm{d}\gamma \\ &= \int_{\Gamma_R^+} \int_0^{\infty} e^{-2s_1 t} \left( \int_0^t \mathscr{L}^{-1} \circ s \mathscr{B} \circ \mathscr{L} \tilde{\boldsymbol{E}}_{\Gamma_R^+}(\tau) \mathrm{d}\tau \right) \cdot \bar{\tilde{\boldsymbol{E}}}_{\Gamma_R^+}(t) \, \mathrm{d}t \mathrm{d}\gamma \\ &= \frac{1}{2\pi} \int_{-\infty}^{\infty} \int_{\Gamma_R^+} \mathscr{B} \circ \mathscr{L} \tilde{\boldsymbol{E}}_{\Gamma_R^+}(s) \cdot \mathscr{L}(\bar{\tilde{\boldsymbol{E}}}_{\Gamma_R^+})(s) \, \mathrm{d}\gamma \mathrm{d}s_2 \\ &= \frac{1}{2\pi} \int_{-\infty}^{\infty} \left\langle \mathscr{B}[\check{\tilde{\boldsymbol{E}}}_{\Gamma_R^+}], \, \check{\tilde{\boldsymbol{E}}}_{\Gamma_R^+} \right\rangle \mathrm{d}s_2 \geq 0. \end{split}$$

The proof is completed by taking  $s_1 \to 0$  in the above inequality.  $\square$ 

THEOREM 4.4. Let  $\mathbf{E} \in \mathbf{H}_S(\text{curl}, \Omega)$  be the solution of (4.11) for any t > 0. Given  $\mathbf{E}_0, \mathbf{E}_1 \in \mathbf{L}^2(\Omega)$  and  $\mathbf{F} \in \mathbf{L}^1(0, T; \mathbf{L}^2(\Omega))$  for any T > 0, there holds

and

$$(4.13) \|\boldsymbol{E}\|_{\boldsymbol{L}^{2}(0,T;\boldsymbol{L}^{2}(\Omega))} \lesssim T^{1/2} \|\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)} + T^{3/2} \|\boldsymbol{E}_{1}\|_{\boldsymbol{L}^{2}(\Omega)} + T^{3/2} \|\boldsymbol{F}\|_{\boldsymbol{L}^{1}(0,T;\boldsymbol{L}^{2}(\Omega))}.$$

*Proof.* Let  $0 < \xi < T$  and define an auxiliary function

(4.14) 
$$\psi(\mathbf{r},t) = \int_{t}^{\xi} \mathbf{E}(\mathbf{r},\tau) d\tau, \quad \mathbf{r} \in \Omega, \quad 0 \le t \le \xi.$$

It is clear that

(4.15) 
$$\psi(\mathbf{r},\xi) = 0, \quad \partial_t \psi(\mathbf{r},t) = -\mathbf{E}(\mathbf{r},t).$$

For any  $\phi(\mathbf{r},t) \in \mathbf{L}^2(0,\xi;\mathbf{L}^2(\Omega))$ , we have

(4.16) 
$$\int_0^{\xi} \phi(\mathbf{r}, t) \cdot \bar{\psi}(\mathbf{r}, t) dt = \int_0^{\xi} \left( \int_0^t \phi(\mathbf{r}, \tau) d\tau \right) \cdot \bar{\mathbf{E}}(\mathbf{r}, t) dt.$$

Indeed, using integration by parts and (4.15), we have

$$\int_{0}^{\xi} \boldsymbol{\phi}(\boldsymbol{r},t) \cdot \bar{\boldsymbol{\psi}}(\boldsymbol{r},t) dt = \int_{0}^{\xi} \left( \boldsymbol{\phi}(\boldsymbol{r},t) \cdot \int_{t}^{\xi} \bar{\boldsymbol{E}}(\boldsymbol{r},\tau) d\tau \right) dt$$

$$= \int_{0}^{\xi} \int_{t}^{\xi} \bar{\boldsymbol{E}}(\boldsymbol{r},\tau) d\tau \cdot d \left( \int_{0}^{t} \boldsymbol{\phi}(\boldsymbol{r},\tau) d\tau \right)$$

$$= \int_{0}^{t} \int_{0}^{t} \boldsymbol{\phi}(\boldsymbol{r},\tau) d\tau \cdot \int_{t}^{\xi} \bar{\boldsymbol{E}}(\boldsymbol{r},\tau) d\tau \Big|_{0}^{\xi} + \int_{0}^{\xi} \left( \int_{0}^{t} \boldsymbol{\phi}(\boldsymbol{r},\tau) d\tau \right) \cdot \bar{\boldsymbol{E}}(\boldsymbol{r},t) dt$$

$$= \int_{0}^{\xi} \left( \int_{0}^{t} \boldsymbol{\phi}(\boldsymbol{r},\tau) d\tau \right) \cdot \bar{\boldsymbol{E}}(\boldsymbol{r},t) dt.$$

Next, we take the test function  $\mathbf{w} = \mathbf{\psi}$  in (4.11) and get

(4.17) 
$$\int_{\Omega} \varepsilon \partial_{t}^{2} \mathbf{E} \cdot \bar{\boldsymbol{\psi}} \, d\mathbf{r} = -\int_{\Omega} \mu^{-1} (\nabla \times \mathbf{E}) \cdot (\nabla \times \bar{\boldsymbol{\psi}}) \, d\mathbf{r} - \int_{\Gamma_{R}^{+}} \mathscr{C}[\mathbf{E}_{\Gamma_{R}^{+}}] \cdot \bar{\boldsymbol{\psi}}_{\Gamma_{R}^{+}} \, d\gamma - \int_{\Omega} \mathbf{F} \cdot \bar{\boldsymbol{\psi}} \, d\mathbf{r}.$$

It follows from (4.15) that

(4.18) 
$$\operatorname{Re} \int_{0}^{\xi} \int_{\Omega} \partial_{t}^{2} \boldsymbol{E} \cdot \bar{\boldsymbol{\psi}} \, d\boldsymbol{r} dt = \operatorname{Re} \int_{\Omega} \int_{0}^{\xi} \left( \partial_{t} (\partial_{t} \boldsymbol{E} \cdot \bar{\boldsymbol{\psi}}) + \partial_{t} \boldsymbol{E} \cdot \bar{\boldsymbol{E}} \right) \, dt d\boldsymbol{r}$$

$$= \operatorname{Re} \int_{\Omega} \left( \left( \partial_{t} \boldsymbol{E} \cdot \bar{\boldsymbol{\psi}} \right) \Big|_{0}^{\xi} + \frac{1}{2} |\boldsymbol{E}|^{2} \Big|_{0}^{\xi} \right) \, d\boldsymbol{r}$$

$$= \frac{1}{2} \|\boldsymbol{E}(\cdot, \xi)\|_{\boldsymbol{L}^{2}(\Omega)}^{2} - \frac{1}{2} \|\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} - \operatorname{Re} \int_{\Omega} \boldsymbol{E}_{1}(\boldsymbol{r}) \cdot \bar{\boldsymbol{\psi}}(\boldsymbol{r}, 0) \, d\boldsymbol{r}.$$

Integrating (4.17) from t = 0 to  $t = \xi$  and taking the real parts yields

$$\frac{\varepsilon}{2} \| \boldsymbol{E}(\cdot, \xi) \|_{\boldsymbol{L}^{2}(\Omega)}^{2} - \frac{\varepsilon}{2} \| \boldsymbol{E}_{0} \|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \frac{1}{2} \int_{\Omega} \mu^{-1} \Big| \int_{0}^{\xi} \nabla \times \boldsymbol{E}(\boldsymbol{r}, t) dt \Big|^{2} d\boldsymbol{r}$$

$$= \varepsilon \operatorname{Re} \int_{\Omega} \boldsymbol{E}_{1}(\boldsymbol{r}) \cdot \bar{\boldsymbol{\psi}}(\boldsymbol{r}, 0) d\boldsymbol{r} - \operatorname{Re} \int_{0}^{\xi} \int_{\Omega} \boldsymbol{F} \cdot \bar{\boldsymbol{\psi}} d\boldsymbol{r}$$

$$- \operatorname{Re} \int_{0}^{\xi} \int_{\Gamma_{R}^{+}} \mathscr{C}[\boldsymbol{E}_{\Gamma_{R}^{+}}] \cdot \bar{\boldsymbol{\psi}}_{\Gamma_{R}^{+}} d\gamma dt.$$
(4.19)

In what follows, we estimate the three terms on the right hand side of (4.19) separately. We derive from the Cauchy-Schwarz inequality that

(4.20) 
$$\operatorname{Re} \int_{\Omega} \boldsymbol{E}_{1}(\boldsymbol{r}) \cdot \bar{\boldsymbol{\psi}}(\boldsymbol{r}, 0) \, d\boldsymbol{r} = \operatorname{Re} \int_{\Omega} \boldsymbol{E}_{1}(\boldsymbol{r}) \cdot \left( \int_{0}^{\xi} \bar{\boldsymbol{E}}(\boldsymbol{r}, t) \, dt \right) d\boldsymbol{r}$$
$$= \operatorname{Re} \int_{0}^{\xi} \int_{\Omega} \boldsymbol{E}_{1}(\boldsymbol{r}) \cdot \bar{\boldsymbol{E}}(\boldsymbol{r}, t) \, d\boldsymbol{r} dt \leq \|\boldsymbol{E}_{1}\|_{\boldsymbol{L}^{2}(\Omega)} \int_{0}^{\xi} \|\boldsymbol{E}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)} \, dt.$$

For  $0 \le t \le \xi \le T$ , we have from (4.16) that

$$\operatorname{Re} \int_{0}^{\xi} \int_{\Omega} \boldsymbol{F} \cdot \bar{\boldsymbol{\psi}} \, \mathrm{d}\boldsymbol{r} \mathrm{d}t = \operatorname{Re} \int_{\Omega} \int_{0}^{\xi} \left( \int_{0}^{t} \boldsymbol{F}(\boldsymbol{r}, \tau) \, \mathrm{d}\tau \right) \cdot \bar{\boldsymbol{E}}(\boldsymbol{r}, t) \, \mathrm{d}t \mathrm{d}\boldsymbol{r}$$

$$= \operatorname{Re} \int_{0}^{\xi} \int_{0}^{t} \int_{\Omega} \boldsymbol{F}(\boldsymbol{r}, \tau) \cdot \bar{\boldsymbol{E}}(\boldsymbol{r}, t) \, \mathrm{d}\boldsymbol{r} \mathrm{d}\tau \, \mathrm{d}t$$

$$\leq \int_{0}^{\xi} \left( \int_{0}^{t} \|\boldsymbol{F}(\cdot, \tau)\|_{\boldsymbol{L}^{2}(\Omega)} \, \mathrm{d}\tau \right) \|\boldsymbol{E}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)} \, \mathrm{d}t$$

$$\leq \int_{0}^{\xi} \left( \int_{0}^{\xi} \|\boldsymbol{F}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)} \, \mathrm{d}t \right) \|\boldsymbol{E}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)} \, \mathrm{d}t$$

$$\leq \left( \int_{0}^{\xi} \|\boldsymbol{F}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)} \, \mathrm{d}t \right) \left( \int_{0}^{\xi} \|\boldsymbol{E}(\cdot, t)\|_{\boldsymbol{L}^{2}(\Omega)} \, \mathrm{d}t \right).$$

$$(4.21)$$

Using Lemma 4.3 and (4.16), we obtain

$$(4.22) \operatorname{Re} \int_{0}^{\xi} \int_{\Gamma} \mathscr{C}(\boldsymbol{E}_{\Gamma}) \cdot \bar{\boldsymbol{\psi}}_{\Gamma} d\gamma dt$$

$$= \operatorname{Re} \int_{\Gamma_{R}^{\pm}} \int_{0}^{\xi} \left( \int_{0}^{t} \mathscr{C}[\boldsymbol{E}_{\Gamma_{R}^{+}}](\boldsymbol{r}, \tau) d\tau \right) \cdot \bar{\boldsymbol{E}}_{\Gamma_{R}^{+}}(\boldsymbol{r}, t) dt d\gamma \geq 0.$$

Substituting (4.20)–(4.22) into (4.19), we have for any  $\xi \in [0,T]$  that

$$\frac{\varepsilon}{2} \| \boldsymbol{E}(\cdot, \xi) \|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \frac{1}{2} \int_{\Omega} \mu^{-1} \Big| \int_{0}^{\xi} \nabla \times \boldsymbol{E}(\boldsymbol{r}, t) dt \Big|^{2} d\boldsymbol{r}$$

$$(4.23) \leq \frac{\varepsilon}{2} \| \boldsymbol{E}_{0} \|_{\boldsymbol{L}^{2}(\Omega)}^{2} + \left( \int_{0}^{\xi} \| \boldsymbol{F}(\cdot, t) \|_{\boldsymbol{L}^{2}(\Omega)} dt + \varepsilon \| \boldsymbol{E}_{1} \|_{\boldsymbol{L}^{2}(\Omega)} \right) \int_{0}^{\xi} \| \boldsymbol{E}(\cdot, t) \|_{\boldsymbol{L}^{2}(\Omega)} dt.$$

Taking the  $L^{\infty}$ -norm with respect to  $\xi$  on both sides of (4.23) yields

$$\|E\|_{L^{\infty}(0,T;L^{2}(\Omega))}^{2} \lesssim \|E_{0}\|_{L^{2}(\Omega)}^{2} + T\left(\|F\|_{L^{1}(0,T;L^{2}(\Omega))} + \|E_{1}\|_{L^{2}(\Omega)}\right) \|E\|_{L^{\infty}(0,T;L^{2}(\Omega))},$$

which gives the estimate (4.12) after applying the Cauchy-Schwarz inequality.

Integrating (4.23) with respect to  $\xi$  from 0 to T and using the Cauchy-Schwarz inequality, we obtain

$$\|\boldsymbol{E}\|_{\boldsymbol{L}^{2}(0,T;\boldsymbol{L}^{2}(\Omega))}^{2} \lesssim T\|\boldsymbol{E}_{0}\|_{\boldsymbol{L}^{2}(\Omega)}^{2} + T^{3/2}\left(\|\boldsymbol{F}\|_{\boldsymbol{L}^{1}(0,T;\boldsymbol{L}^{2}(\Omega))} + \|\boldsymbol{E}_{1}\|_{\boldsymbol{L}^{2}(\Omega)}\right)\|\boldsymbol{E}\|_{\boldsymbol{L}^{2}(0,T;\boldsymbol{L}^{2}(\Omega))},$$

which implies the estimate (4.13) by using the Cauchy-Schwarz inequality again.  $\square$ 

In Theorem 4.4, it is required that  $E_0, E_1 \in L^2(\Omega)$ , and  $F \in L^1(0, T; L^2(\Omega))$ , which can be satisfied if the data satisfy

$$E_0 \in L^2(\Omega), \quad H_0 \in H(\text{curl}, \Omega), \quad J \in H^1(0, T; L^2(\Omega)).$$

It is important to point out that the estimates in Theorem 4.2 were derived from a usual energy method, while the results in Theorem 4.4 were obtained by using different test functions (cf. (4.14)).

**5. Concluding remarks.** In this paper, we studied the time-domain Maxwell equations in a three-dimensional cavity. The open cavity scattering problem was reduced to an initial-boundary value problem in a bounded domain by using the exact transparent boundary condition. The reduced problem was shown to have a unique solution and its stability was also presented. The proofs were based on the examination of the time-harmonic Maxwell equations with complex wavenumber and the abstract inversion theorem of the Laplace transform. Moreover, by directly considering the variational problem of the time-domain Maxwell equations, an *a priori* estimate was derived with an explicit dependence on time for the electric field. Computationally, the variational approach leads naturally to a class of finite element methods. As a time-dependent problem, a fast and accurate marching technique shall be developed to deal with the temporal convolution in the transparent boundary condition. We will report the work on its numerical analysis and computation in a forthcoming paper.

#### Appendix A. Laplace transform.

For any  $s = s_1 + \mathrm{i} s_2$  with  $s_1, s_2 \in \mathbb{R}, s_1 > 0$ , define by  $\boldsymbol{u}(s)$  the Laplace transform of the vector field  $\boldsymbol{u}(t)$ , i.e.,

$$\check{\boldsymbol{u}}(s) = \mathscr{L}(\boldsymbol{u})(s) = \int_0^\infty e^{-st} \boldsymbol{u}(t) \, \mathrm{d}t.$$

It can be verified from the integration by parts that

(A.1) 
$$\int_0^t \boldsymbol{u}(\tau) \, d\tau = \mathcal{L}^{-1} \left( s^{-1} \boldsymbol{\breve{u}}(s) \right),$$

where  $\mathcal{L}^{-1}$  is the inverse Laplace transform. One verifies from the formula of the inverse Laplace transform that

(A.2) 
$$\mathbf{u}(t) = \mathscr{F}^{-1}(e^{s_1 t} \mathscr{L}(\mathbf{u})(s_1 + \mathrm{i} s_2)),$$

where  $\mathscr{F}^{-1}$  denotes the inverse Fourier transform with respect to  $s_2$ .

Recall the Plancherel or Parseval identity for the Laplace tranform (cf., [13, (2.46)]).

LEMMA A.1. If  $\breve{\mathbf{u}} = \mathcal{L}(\mathbf{u})$  and  $\breve{\mathbf{v}} = \mathcal{L}(\mathbf{v})$ , then

$$\frac{1}{2\pi} \int_{-\infty}^{\infty} \check{\boldsymbol{u}}(s) \cdot \check{\boldsymbol{v}}(s) ds_2 = \int_{0}^{\infty} e^{-2s_1 t} \boldsymbol{u}(t) \cdot \boldsymbol{v}(t) dt,$$

for all  $s_1 > \lambda$ , where  $\lambda$  is the absissa of convergence for the Laplace transform of  $\boldsymbol{u}$  and  $\boldsymbol{v}$ .

The following theorem (cf., [33, Thm 43.1]) is an analog of the Paley-Wiener-Schwarz theorem for the Fourier transform of the distributions with compact support in the case of Laplace transform.

LEMMA A.2. Let  $\check{\boldsymbol{w}}(s)$  denote a holomorphic function in the half plane  $s_1 > \sigma_0$ , valued in the Banach space  $\mathbb{E}$ . The following statements are equivalent:

- 1. there is a distribution  $\mathbf{w} \in \mathcal{D}'_{+}(\mathbb{E})$  whose Laplace transform is equal to  $\check{\mathbf{w}}(s)$ ;
- 2. there is a  $\sigma_1$  with  $\sigma_0 \leq \sigma_1 < \infty$  and an interger  $m \geq 0$  such that for all complex numbers s with  $s_1 > \sigma_1$ , it holds  $\|\check{\boldsymbol{w}}(s)\|_{\mathbb{E}} \lesssim (1+|s|)^m$ ,

where  $\mathcal{D}'_{+}(\mathbb{E})$  is the space of distributions on the real line which vanish identically in the open negative half line.

## Appendix B. Spherical harmonics on hemisphere.

The spherical coordinates  $(\rho, \theta, \varphi)$  are related to the Cartesian coordinates  $r = (x_1, x_2, x_3)$  by  $x_1 = \rho \sin \theta \cos \varphi$ ,  $x_2 = \rho \sin \theta \sin \varphi$ ,  $x_3 = \rho \cos \theta$ , with the local orthonormal basis  $\{e_\rho, e_\theta, e_\varphi\}$ :

$$\begin{aligned} & \boldsymbol{e}_{\rho} = (\sin\theta\cos\varphi, \sin\theta\sin\varphi, \cos\theta), \\ & \boldsymbol{e}_{\theta} = (\cos\theta\cos\varphi, \cos\theta\sin\varphi, -\sin\theta), \\ & \boldsymbol{e}_{\varphi} = (-\sin\varphi, \cos\varphi, 0), \end{aligned}$$

where  $\theta$  and  $\varphi$  are the Euler angles. Let  $\Gamma = \{r : \rho = 1\}$ ,  $\Gamma^+ = \{r : \rho = 1, x_3 \ge 0\}$ ,  $\Gamma^- = \{r : \rho = 1, x_3 \le 0\}$  be the unit sphere, upper unit hemisphere, and lower unit hemisphere, respectively. Denote by  $\Gamma_R = \{r : \rho = R\}$ ,  $\Gamma_R^+ = \{r : \rho = R, x_3 \ge 0\}$ ,  $\Gamma_R^- = \{r : \rho = R, x_3 \le 0\}$  the whole sphere, upper hemisphere, and lower hemisphere with radius R, respectively.

Let  $\{Y_n^m(\theta,\varphi), |m| \le n, n = 0, 1, 2, ...\}$  be an orthonormal sequence of spherical harmonics of order n on the unit sphere  $\Gamma$  that satisfies

(B.1) 
$$\Delta_{\Gamma} Y_n^m + n(n+1) Y_n^m = 0,$$

where

$$\Delta_{\Gamma} = \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \varphi^2}$$

is the Laplace-Beltrami operator on  $\Gamma$ . Explicitly, the spherical harmonics of order n is written as

$$Y_n^m(\theta,\varphi) = \sqrt{\frac{2n+1}{4\pi} \frac{(n-|m|)!}{(n+|m|)!}} P_n^{|m|}(\cos\theta) e^{\mathrm{i}m\varphi},$$

where the associated Legendre functions are

$$P_n^m(t) := \sqrt{(1-t^2)^m} \frac{\mathrm{d}^m P_n(t)}{\mathrm{d}t^m}, \quad m = 0, 1, \dots, n.$$

Here  $P_n$  is the Legendre polynomial of degree n, which is an even function if n is even and an odd function if n is odd.

Define a sequence of rescaled spherical harmonics of order n:

$$(\mathrm{B.2}) \hspace{3.1em} X_n^m(\theta,\varphi) = \frac{\sqrt{2}}{R} \, Y_n^m(\theta,\varphi),$$

which forms a complete orthonormal system in  $L^2(\Gamma_R^+)$  for  $|m| \le n$ , m+n = odd,  $n \in \mathbb{N}$  (cf., [29, Lemma 3.1]).

Denote by  $L^2(\Gamma_R^+)$  the complex square integrable functions on the hemisphere  $\Gamma_R^+$ . For convenience, we take the following notation for double summations:

$$\sum_{|m| \le n} w_n^m := \sum_{n=1}^{\infty} \sum_{m=-n}^n w_n^m,$$

$$\sum_{|m| \le n}^{\text{odd}} w_n^m := \sum_{n=1}^{\infty} \sum_{\substack{m=-n \\ m+n=\text{odd}}}^n w_n^m,$$

$$\sum_{|m| \le n}^{\text{even}} w_n^m := \sum_{n=1}^{\infty} \sum_{\substack{m=-n \\ m+n=\text{even}}}^n w_n^m.$$

To describe vector wave functions on the hemisphere, we introduce some boundary differential operators. For a smooth scalar function w defined on  $\Gamma_R^+$ , let

$$\nabla_{\Gamma} w = \frac{\partial w}{\partial \theta} \, \boldsymbol{e}_{\theta} + \frac{1}{\sin \theta} \frac{\partial w}{\partial \varphi} \, \boldsymbol{e}_{\varphi}$$

be the tangential gradient on  $\Gamma_R^+$ . The surface vector curl is defined by

$$\boldsymbol{c}url_{\Gamma}w = \nabla_{\Gamma}w \times \boldsymbol{e}_{o}.$$

Denote by  $\operatorname{div}_{\Gamma}$  and  $\operatorname{curl}_{\Gamma}$  the surface divergence and the surface scalar curl, respectively. For a smooth vector function  $\boldsymbol{w}$  tangential to  $\Gamma_R^+$ , it can be represented by its coordinates in the local orthonormal basis:

$$\boldsymbol{w} = w_{\theta} \, \boldsymbol{e}_{\theta} + w_{\varphi} \, \boldsymbol{e}_{\varphi},$$

where

$$w_{\theta} = \boldsymbol{w} \cdot \boldsymbol{e}_{\theta}$$
 and  $w_{\varphi} = \boldsymbol{w} \cdot \boldsymbol{e}_{\varphi}$ .

The surface divergence and the surface scalar curl can be defined as

$$\operatorname{div}_{\Gamma} \boldsymbol{w} = \frac{1}{\sin \theta} \left[ \frac{\partial}{\partial \theta} (w_{\theta} \sin \theta) + \frac{\partial w_{\varphi}}{\partial \varphi} \right],$$
$$\operatorname{curl}_{\Gamma} \boldsymbol{w} = \frac{1}{\sin \theta} \left[ \frac{\partial}{\partial \theta} (w_{\varphi} \sin \theta) - \frac{\partial w_{\theta}}{\partial \varphi} \right].$$

It is known (cf., [32]) that these boundary differential operators satisfy

(B.3) 
$$\Delta_{\Gamma} = \operatorname{div}_{\Gamma} \nabla_{\Gamma} = -\operatorname{curl}_{\Gamma} \operatorname{curl}_{\Gamma} \quad \text{and} \quad \operatorname{curl}_{\Gamma} \nabla_{\Gamma} = \operatorname{div}_{\Gamma} \operatorname{curl}_{\Gamma} = 0.$$

It is also know (cf., [14, Thm 6.23]) that an orthonormal basis for  $L_{\rm t}^2(\Gamma_R) = \{ \boldsymbol{w} \in L^2(\Gamma_R) : \boldsymbol{e}_{\rho} \cdot \boldsymbol{w} = 0 \}$ , the tangential fields on  $\Gamma_R$ , consists of functions of the form

$$\boldsymbol{U}_{n}^{m}(\theta,\varphi) = \frac{1}{R\sqrt{n(n+1)}} \nabla_{\Gamma} Y_{n}^{m}(\theta,\varphi)$$

and

$$\boldsymbol{V}_{n}^{m}(\theta,\varphi) = \boldsymbol{e}_{\rho} \times \boldsymbol{U}_{n}^{m}(\theta,\varphi) = -\frac{1}{R\sqrt{n(n+1)}}\boldsymbol{c}url_{\Gamma}Y_{n}^{m}$$

for  $|m| \leq n$ ,  $n \in \mathbb{N}$ . It follows from (B.1) and (B.3) that

$$\operatorname{div}_{\Gamma} \boldsymbol{U}_{n}^{m} = -\frac{\sqrt{n(n+1)}}{R} Y_{n}^{m}, \quad \operatorname{curl}_{\Gamma} \boldsymbol{V}_{n}^{m} = -\frac{\sqrt{n(n+1)}}{R} Y_{n}^{m},$$

and

$$\operatorname{curl}_{\Gamma} \boldsymbol{U}_{n}^{m} = \operatorname{div}_{\Gamma} \boldsymbol{V}_{n}^{m} = 0.$$

Define two sequences of tangential fields

(B.4) 
$$\boldsymbol{X}_{n}^{m}(\theta,\varphi) = \frac{1}{\sqrt{n(n+1)}} \nabla_{\Gamma} X_{n}^{m}(\theta,\varphi) = \sqrt{2} \boldsymbol{U}_{n}^{m}(\theta,\varphi)$$

and

$$\boldsymbol{Y}_{n}^{m}(\theta,\varphi) = \boldsymbol{e}_{\rho} \times \boldsymbol{X}_{n}^{m}(\theta,\varphi) = \sqrt{2} \, \boldsymbol{V}_{n}^{m}(\theta,\varphi).$$

Using the definition of the tangential gradient, and noticing that  $e_{\theta} \times e_{\varphi} = e_{\rho}$ ,  $e_{\varphi} \times e_{\rho} = e_{\theta}$ ,  $e_{\rho} \times e_{\theta} = e_{\varphi}$ , we get

$$e_{\theta} \times \boldsymbol{X}_{n}^{m}(\frac{\pi}{2}, \varphi) = 0 \text{ for } |m| \leq n, m+n = \text{odd}, n \in \mathbb{N},$$

and

$$e_{\theta} \times \boldsymbol{Y}_{n}^{m}(\frac{\pi}{2}, \varphi) = 0 \text{ for } |m| \leq n, m+n = \text{even}, n \in \mathbb{N}.$$

Define a subspace of complex square integrable tangential fields functions on the hemisphere  $\Gamma_R^+$ :

$$L_{\mathrm{t}}^{2}(\Gamma_{R}^{+}) = \{ w \in L^{2}(\Gamma_{R}^{+}) : e_{\rho} \cdot w = 0 \}.$$

It is shown (cf., [29, Lemma 3.2]) that the vector spherical harmonics  $\{\boldsymbol{X}_n^m: m+n=\text{odd}\}$  and  $\{\boldsymbol{Y}_n^m: m+n=\text{even}\}$  for  $|m|\leq n, n\in\mathbb{N}$  form a complete orthonormal system in  $\boldsymbol{L}_{\mathrm{t}}^2(\Gamma_R^+)$ .

### Appendix C. Hankel functions.

For  $\nu \in \mathbb{R}$ , the two Hankel functions  $H_{\nu}^{(1)}(z)$  and  $H_{\nu}^{(2)}(z)$  where  $z \in \mathbb{C}$ , are two fundamental solutions of the Bessel equation for functions of order  $\nu$ :.

$$z^{2} \frac{\mathrm{d}^{2} u}{\mathrm{d}z^{2}} + z \frac{\mathrm{d}u}{\mathrm{d}z} + (z^{2} - \nu^{2})u = 0.$$

Recall the Bessel functions of imaginary argument  $K_{\nu}(z)$ , also called the modified Bessel functions, which is the solution of the differential equation

$$z^{2} \frac{\mathrm{d}^{2} u}{\mathrm{d}z^{2}} + z \frac{\mathrm{d}u}{\mathrm{d}z} - (z^{2} + \nu^{2})u = 0.$$

It is connected with  $H_{\nu}^{(1)}(z)$  through the relation

(C.1) 
$$K_{\nu}(z) = \frac{1}{2}\pi i e^{\frac{1}{2}\nu\pi i} H_{\nu}^{(1)}(iz).$$

It is known (cf., [38, p. 511]) that  $K_{\nu}(z)$  has no zeros if  $|\arg z| \leq \frac{1}{2}\pi$ , which implies from (C.1) that  $H_{\nu}^{(1)}(z)$  has no zeros when  $\text{Im}z \leq 0$ .

The spherical Hankel function  $h_n^{(1)}(z)$  can also be defined by the Hankel function of half integer order:

(C.2) 
$$h_n^{(1)}(z) = \sqrt{\frac{\pi}{2z}} H_{n+\frac{1}{2}}^{(1)}(z).$$

Combining (C.1) and (C.2) yields

$$h_n^{(1)}(z) = -\sqrt{\frac{2\pi}{z}} i e^{-\frac{1}{2}(n+\frac{1}{2})\pi i} K_{n+\frac{1}{2}}(-iz),$$

which implies that  $h_n^{(1)}(z)$  has no zeros when  $\text{Im} z \geq 0$ .

The following two lemmas on the spherical Hankel functions for the complex number are proved in [11].

LEMMA C.1. Let  $R > 0, n \in \mathbb{Z}, s = s_1 + is_2$  with  $s_1 > 0$ . It holds that

$$\operatorname{Re}\left\{1 + r_n^{(1)}(\mathrm{i}sR)\right\} < 0,$$

where  $r_n^{(1)}(z) = z(h_n^{(1)})'(z)/h_n^{(1)}(z)$ . Lemma C.2. Let  $R>0, n\in\mathbb{Z}, s=s_1+\mathrm{i} s_2$  with  $s_1>0$ . Then we have

$$\operatorname{Im}\left\{s_2(1+r_n^{(1)}(\mathrm{i}sR))\right\} \le 0.$$

Combining the result in Lemma C.1 for small value of n and the proof in [14] (cf. [25, Lemma 3.1]) for large value of n, we may obtain the following estimate.

LEMMA C.3. Let  $R > 0, n \in \mathbb{Z}, s = s_1 + is_2$  with  $s_1 > 0$ . There exist positive constants  $C_1$  and  $C_2$  such that

$$C_1 n \le |1 + r_n^{(1)}(isR)| \le C_2 n.$$

### Appendix D. Functional spaces.

Denote by  $H^s(\Gamma_R^+)$  the Sobolev space, the completion of  $C_0^{\infty}(\Gamma_R^+)$  in the norm  $\|\cdot\|_{H^s(\Gamma_R^+)}$  characterized by

$$||w||_{H^s(\Gamma_R^+)}^2 = \sum_{|m| \le n}^{\text{odd}} (1 + n(n+1))^s |w_n^m|^2,$$

where

$$w(\theta, \varphi) = \sum_{|m| \le n}^{\text{odd}} w_n^m X_n^m(\theta, \varphi).$$

Introduce three tangential trace spaces:

$$\boldsymbol{H}_{t}^{s}(\Gamma_{R}^{+}) = \left\{ \boldsymbol{w} \in (H^{s}(\Gamma_{R}^{+}))^{3}, \ \boldsymbol{e}_{\rho} \cdot \boldsymbol{w} = 0, \ \boldsymbol{e}_{\theta} \times \boldsymbol{w}(\pi/2, \varphi) = 0 \right\}, \\
\boldsymbol{H}^{-1/2}(\operatorname{curl}, \Gamma_{R}^{+}) = \left\{ \boldsymbol{w} \in \boldsymbol{H}_{t}^{-1/2}(\Gamma_{R}^{+}), \ \operatorname{curl}_{\Gamma} \boldsymbol{w} \in H^{-1/2}(\Gamma_{R}^{+}) \right\}, \\
\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_{R}^{+}) = \left\{ \boldsymbol{w} \in \boldsymbol{H}_{t}^{-1/2}(\Gamma_{R}^{+}), \ \operatorname{div}_{\Gamma} \boldsymbol{w} \in H^{-1/2}(\Gamma_{R}^{+}) \right\}.$$

For any tangential field  $\boldsymbol{w} \in \boldsymbol{H}_{\mathrm{t}}^{s}(\Gamma_{R}^{+})$ , it can be represented in the series expansion

$$\boldsymbol{w} = \sum_{|m| \le n}^{\text{odd}} w_{1n}^m \boldsymbol{X}_n^m(\theta, \varphi) + \sum_{|m| \le n}^{\text{even}} w_{2n}^m \boldsymbol{Y}_n^m(\theta, \varphi).$$

Using the series coefficients, the norm of the space  $H_{\rm t}^s(\Gamma_R^+)$  can be characterized by

$$\|\boldsymbol{w}\|_{\boldsymbol{H}_{t}^{s}(\Gamma_{R}^{+})}^{2} = \sum_{|m| \leq n}^{\text{odd}} (1 + n(n+1))^{s} |w_{1n}^{m}|^{2} + \sum_{|m| \leq n}^{\text{even}} (1 + n(n+1))^{s} |w_{2n}^{m}|^{2};$$

the norm of the space  $\boldsymbol{H}^{-1/2}(\operatorname{curl},\,\Gamma_R^+)$  can be characterized by

$$\|\boldsymbol{w}\|_{\boldsymbol{H}^{-1/2}(\text{curl},\,\Gamma_R^+)}^2 = \sum_{|m| \le n}^{\text{odd}} \frac{1}{\sqrt{1 + n(n+1)}} \, |w_{1n}^m|^2 + \sum_{|m| \le n}^{\text{even}} \sqrt{1 + n(n+1)} \, |w_{2n}^m|^2;$$

and the norm of the space  $H^{-1/2}(\text{div}, \Gamma_R^+)$  can be characterized by

$$\|\boldsymbol{w}\|_{\boldsymbol{H}^{-1/2}(\operatorname{div}, \Gamma_R^+)}^2 = \sum_{|m| \le n}^{\operatorname{odd}} \sqrt{1 + n(n+1)} |w_{1n}^m|^2 + \sum_{|m| \le n}^{\operatorname{even}} \frac{1}{\sqrt{1 + n(n+1)}} |w_{2n}^m|^2.$$

Define a dual pairing by

$$\langle \boldsymbol{u}, \boldsymbol{v} \rangle = \int_{\Gamma_R^+} \boldsymbol{u} \cdot \bar{\boldsymbol{v}} \, \mathrm{d}\gamma = \sum_{|m| < n}^{\mathrm{odd}} u_{1n}^m \, \bar{v}_{1n}^m + \sum_{|m| < n}^{\mathrm{even}} u_{2n}^m \, \bar{v}_{2n}^m,$$

where

$$\boldsymbol{u} = \sum_{|m| \le n}^{\text{odd}} u_{1n}^m \boldsymbol{X}_n^m + \sum_{|m| \le n}^{\text{even}} u_{2n}^m \boldsymbol{Y}_n^m \quad \text{and} \quad \boldsymbol{v} = \sum_{|m| \le n}^{\text{odd}} v_{1n}^m \boldsymbol{X}_n^m + \sum_{|m| \le n}^{\text{even}} v_{2n}^m \boldsymbol{Y}_n^m.$$

Introduce three functional spaces

$$\begin{aligned} & \boldsymbol{H}(\operatorname{curl},\,\Omega) = \big\{\boldsymbol{u} \in \boldsymbol{L}^2(\Omega), \,\, \nabla \times \boldsymbol{E} \in \boldsymbol{L}^2(\Omega)\big\}, \\ & \boldsymbol{H}_S(\operatorname{curl},\,\Omega) = \big\{\boldsymbol{u} \in \boldsymbol{H}(\operatorname{curl},\,\Omega), \,\, \boldsymbol{n} \times \boldsymbol{u} = 0 \,\,\text{on} \,\, S\big\}, \\ & \boldsymbol{H}_0(\operatorname{curl},\,\Omega) = \big\{\boldsymbol{u} \in \boldsymbol{H}(\operatorname{curl},\,\Omega), \,\, \boldsymbol{n} \times \boldsymbol{u} = 0 \,\,\text{on} \,\, S \cup \Gamma_R^+\big\}, \end{aligned}$$

which are Sobolev spaces with the norm:

$$\|\boldsymbol{u}\|_{\boldsymbol{H}(\mathrm{curl},\,\Omega)} = \left(\|\boldsymbol{u}\|_{\boldsymbol{L}^2(\Omega)}^2 + \|
abla imes \boldsymbol{u}\|_{\boldsymbol{L}^2(\Omega)}^2
ight)^{1/2}.$$

Given a vector field  $\boldsymbol{u}$  on  $\Gamma_R^+$ , denote by

$$oldsymbol{u}_{\Gamma_{B}^{+}}=-oldsymbol{e}_{
ho} imes(oldsymbol{e}_{
ho} imesoldsymbol{u})$$

the tangential component of  $\boldsymbol{u}$  on  $\Gamma_R^+$ , which satisfies the following trace estimate (cf. [29, Lemma 3.3]).

LEMMA D.1. For any  $\mathbf{u} \in \mathbf{H}_S(\text{curl}, \Omega)$ , there holds

$$\|\boldsymbol{u}_{\Gamma_R^+}\|_{\boldsymbol{H}^{-1/2}(\operatorname{curl},\,\Gamma_R^+)} \lesssim \|\boldsymbol{u}\|_{\boldsymbol{H}(\operatorname{curl},\,\Omega)}.$$

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